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**Key words.** meteorites, meteors, meteoroids – minor planets, asteroids: general – minor planets, asteroids: individual: 24-themis – minor planets, asteroids: individual: 1-ceres – errata, addenda

An error occurred during the production process. Figure 1 was published twice. The corrected Figs. 1 and 4 are published below.

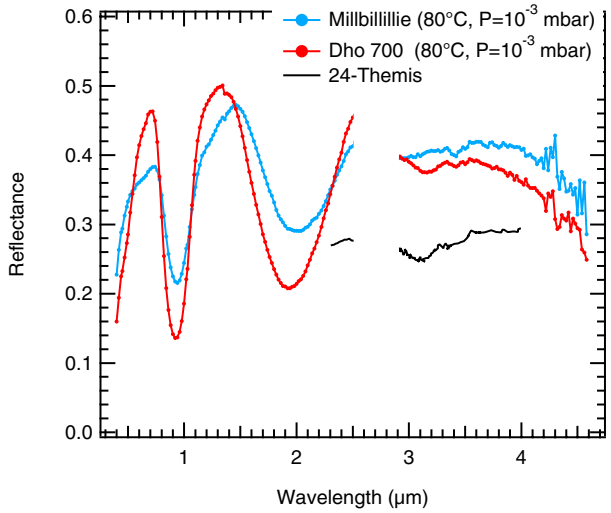


**Fig. 1.** NIR spectra of a eucrite (Millbillillie) and a diogenite powder (Dhofar700) showing the absence of significant $H_2O$- related 3-$\mu$m band upon heating to 80 °C and in a moderate vacuum ($P = 10^{-3}$ mbar). Note the presence of a 3.1 $\mu$m band in the spectrum of Dhofar 700. From its shape and position, this feature may be caused by a small amount of goethite produced by terrestrial weathering. The small absorption at 3.4 $\mu$m is likely due to the presence of organic contamination.

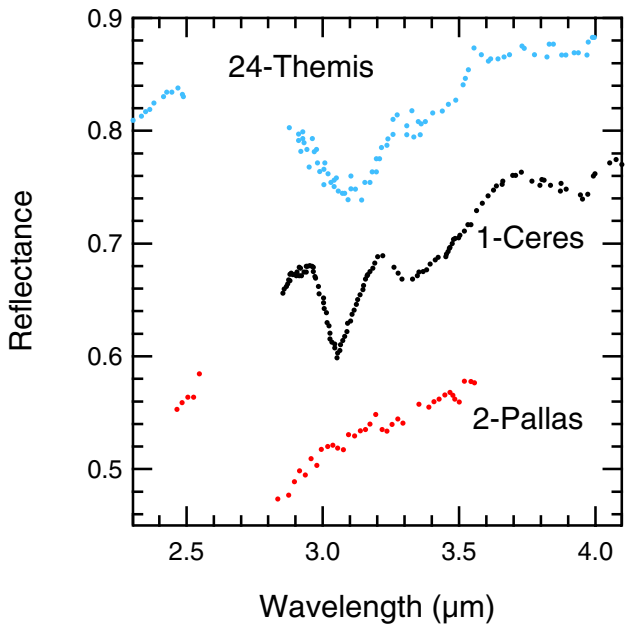


**Fig. 4.** Different types of 3-$\mu$m band as observed on asteroids. 1-Ceres spectrum is from Milliken & Rivkin (2009), 2-Pallas spectrum is from Jones et al. (1990), and 24-Themis data are from Rivkin & Emery (2010). Spectra have been offset and rescaled for clarity.

## References

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