The SuperCam Remote Sensing Instrument Suite for Mars 2020


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THE SUPERCAM REMOTE SENSING INSTRUMENT SUITE FOR MARS 2020

Overview: The Mars 2020 rover, essentially a structural twin of MSL, is being built to a) characterize the geology and history of a new landing site on Mars, b) find and characterize ancient habitable environments, c) cache samples for eventual return to Earth, and d) demonstrate in situ production of oxygen needed for human exploration. Remote-sensing instrumentation is needed to support the first three of these goals [1]. The SuperCam instrument meets these needs with a range of instrumentation including the highest-resolution remote imaging on the rover, two different techniques for determining mineralogy, and one technique to provide elemental compositions. All of these techniques are co-bore sighted, providing rapid comprehensive characterization. In addition, for targets within 7 meters of the rover the laser shock waves brush away the dust, providing cleaner surfaces for analysis. SuperCam will use an advanced version of the AEGIS robotic target selection software.

Instrument Architecture: SuperCam’s overall design (Fig. 1) is strongly patterned after ChemCam, being divided into a Mast Unit (MU) and Body Unit (BU) with an optical fiber and electrical cables connecting the units. The MU includes a 110 mm telescope, a Nd:YAG laser (1064 & 532 nm beams), an IR spectrometer, the Remote Micro-Imager (RMI) camera, and associated electronics. The Body Unit consists of two reflection spectrometers of ChemCam heritage plus an all-purpose transmission spectrometer for laser-induced breakdown (LIBS), Raman, and VIS spectroscopy. The MU is built in France, while the BU is built at LANL. A set of rover calibration targets are being assembled in Spain.

Visible and Infrared (VISIR) Spectroscopy: Mineralogy detection is performed by both Raman and VISIR reflectance spectroscopy. The VISIR spectral range extends from 0.4 to 2.6 µm with two gaps, one at 0.47-0.54, and the other at 0.86-1.3 µm. Spectrometers in the BU cover the range from 0.4-0.86 µm at high spectral resolution, while the 1.3-2.6 µm range is covered in 248 channels by a wavelength-scanning AOTF spectrometer in the MU [2]. The spectral resolution of the latter is 30 cm⁻¹ with sub-sampling to 15 cm⁻¹. Fig. 2 shows a spectrum from the prototype spectrometer. A minimum signal-to-noise ratio of 45 is expected to be obtained within the required exposure times. Spatial resolution will be 0.8-1.1 mrad.
mineralogy. An example spectrum from the prototype SuperCam Raman spectrometer [5] is shown in Fig. 3. The flight unit will have improved spectral resolution of 10 cm⁻¹.

**LIBS:** The ChemCam instrument has currently returned elemental compositions from more than 7,000 observation points in Gale crater (Fig. 4), demonstrating the important contribution of LIBS to planetary exploration. SuperCam will have a very similar LIBS capability to that of ChemCam, with observations to 7 m distance, and with spatial resolution of 300-500 µm [6]. The LIBS will have a similar depth profiling capability, which on ChemCam has discovered thin layers of alteration (Li and other alkali elements [7] and Mn oxides [8].)

**RMI:** The RMI imager on SuperCam uses a CMOS detector with a Bayer color filter to provide the highest resolution remote images (e.g., 0.040 mrad in a ~20 mrad field of view, similar to ChemCam) in color [9]. These images use a high-dynamic-range compression to increase the intensity range using several exposures. The RMI detector and front-end electronics are encapsulated in a compact cube. A simulated image is shown in Fig. 5.

**Current Status:** SuperCam passed its preliminary design review in October and in 2016 it is proceeding with assembly and testing of the development unit, which will verify the Raman timing requirements and signal-to-noise ratios.

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