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# Simulation of the diurnal variations of the oxygen isotope anomaly $(\Delta^{17}O)$ of reactive atmospheric species

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Abstract. The isotope anomaly ( $\Delta^{17}$ O) of secondary atmospheric species such as nitrate (NO<sub>3</sub><sup>-</sup>) or hydrogen peroxide (H<sub>2</sub>O<sub>2</sub>) has potential to provide useful constrains on their formation pathways. Indeed, the  $\Delta^{17}$ O of their precursors (NO<sub>x</sub>, HO<sub>x</sub> etc.) differs and depends on their interactions with ozone, which is the main source of non-zero  $\Delta^{17}$ O in the atmosphere. Interpreting variations of  $\Delta^{17}$ O in secondary species requires an in-depth understanding of the  $\Delta^{17}$ O of their precursors taking into account non-linear chemical regimes operating under various environmental settings.

This article reviews and illustrates a series of basic concepts relevant to the propagation of the  $\Delta^{17}O$  of ozone to other reactive or secondary atmospheric species within a photochemical box model. We present results from numerical simulations carried out using the atmospheric chemistry box model CAABA/MECCA to explicitly compute the diurnal variations of the isotope anomaly of short-lived species such as NO<sub>x</sub> and HO<sub>x</sub>. Using a simplified but realistic tropospheric gas-phase chemistry mechanism,  $\Delta^{17}O$  was propagated from ozone to other species (NO, NO<sub>2</sub>, OH, HO<sub>2</sub>, RO<sub>2</sub>, NO<sub>3</sub>, N<sub>2</sub>O<sub>5</sub>, HONO, HNO<sub>3</sub>, HNO<sub>4</sub>, H<sub>2</sub>O<sub>2</sub>) according to the mass-balance equations, through the implementation of various sets of hypotheses pertaining to the transfer of  $\Delta^{17}O$  during chemical reactions.

The model results confirm that diurnal variations in  $\Delta^{17}O$  of NO<sub>x</sub> predicted by the photochemical steady-state relationship during the day match those from the explicit treatment, but not at night. Indeed, the  $\Delta^{17}O$  of NO<sub>x</sub> is "frozen" at night



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as soon as the photolytical lifetime of NO<sub>x</sub> drops below ca. 10 min. We introduce and quantify the diurnally-integrated isotopic signature (DIIS) of sources of atmospheric nitrate and H<sub>2</sub>O<sub>2</sub>, which is of particular relevance to larger-scale simulations of  $\Delta^{17}$ O where high computational costs cannot be afforded.

#### 1 Introduction

Unraveling chemical mechanisms at play in the atmosphere requires finding creative ways to test the predictions of models which describe them. Most studies to date have relied on concentration measurements to validate model results. Over the past decades alternative isotopic approaches have demonstrated great capabilities in providing concentrationindependent information relevant to atmospheric processes (Laj et al., 2009; Monks et al., 2009). Of particular interest is the development of measurements of the isotope anomaly  $(\Delta^{17}O)$  of oxygen-bearing species (Thiemens, 2006).  $\Delta^{17}O$ is defined as  $\delta^{17}O - 0.52 \times \delta^{18}O$ , with  $\delta^{x}O = R^{x}/R_{VSMOW}^{x} - 1$ (x = 17 or 18) where  $R^x$  refers to the  ${}^xO/{}^{16}O$  elemental ratio in the species of interest and in Vienna Standard Mean Ocean Water (VSMOW), taken as a reference. Ozone  $(O_3)$ possesses a distinctive isotope anomaly inherited from nonmass dependent fractionation NMDF during its formation in the atmosphere (Marcus, 2008).

In contrast to conventional isotopic ratios which are strongly affected by isotopic fractionation,  $\Delta^{17}$ O is fairly insensitive to mass-dependent fractionation. The vast majority of chemical reactions induce mass-dependent fractionation, which in general do not strongly modify  $\Delta^{17}$ O. It can thus be reasonably assumed that  $\Delta^{17}$ O is transferred as is during oxidation reactions in the atmosphere, although this consideration is sometimes challenged when dealing with species possessing small  $\Delta^{17}$ O values (Kaiser et al., 2004). As a result,  $\Delta^{17}$ O of a given species generally simply reflects the fractional importance in its elemental composition of oxygen atoms inherited directly or indirectly from ozone. This behavior has opened large possibilities to explore atmospheric oxidation mechanisms using  $\Delta^{17}$ O signatures (Lyons, 2001; Michalski et al., 2003; Thiemens, 2006; Savarino et al., 2000; Savarino and Morin, 2011).

One area of intense research on the interpretation of  $\Delta^{17}$ O signatures is the case of inorganic atmospheric nitrate (HNO<sub>3</sub> + particulate  $NO_3^-$ ), referred to as atmospheric nitrate  $(NO_3^-)$ below. Indeed, atmospheric nitrate is the final oxidation product of nitrogen oxides  $(NO_x = NO + NO_2)$ , which are of primary importance for air-quality (Jacob, 1999; Finlayson-Pitts and Pitts, 2000; Brown et al., 2006). The development of sensitive methods to analyze the oxygen isotopic composition of nitrate (Michalski et al., 2002; Kaiser et al., 2007) makes it possible to obtain  $\Delta^{17}$ O of atmospheric nitrate at weekly to sub-daily timescales in most environments. This has been used in the recent past to study the seasonal variations in NO<sub>x</sub> oxidation pathways in mid-latitudes (Michalski et al., 2003; Tsunogai et al., 2010) and polar (Morin et al., 2008, 2009; Kunasek et al., 2008) regions, the nature of the sources of atmospheric nitrate in the Antarctic lower atmosphere (Savarino et al., 2007; McCabe et al., 2007; Frey et al., 2009), and more recently the global-scale variations in  $NO_x$ sink reactions (Alexander et al., 2009).  $\Delta^{17}$ O of nitrate has also been used to identify long-term changes in the oxidative properties of the Earth atmosphere, from centennial (Alexander et al., 2004) to millenial (Erbland et al., 2009) time scales.

While including the isotopic composition of long-lived tracers (e.g. CO2, N2O etc.) into global biogeochemical models of the carbon and nitrogen cycle has proved extremely successful (e.g., Hoag et al., 2005), embedding the  $\Delta^{17}$ O of short-lived reactive compounds into atmospheric photochemical models has only recently gained increased attention (Lyons, 2001; Michalski et al., 2003; Zahn et al., 2006; Dominguez et al., 2009; Gromov et al., 2010; Michalski and Xu, 2010). Current hope within the "atmospheric geochemistry community" is that  $\Delta^{17}$ O data can help solve atmospheric chemistry issues such as ascertaining the relative role of heterogeneous reactions in NO<sub>x</sub> sink mechanisms (i.e. what is the exact role of N<sub>2</sub>O<sub>5</sub> hydrolysis; Brown et al., 2006). However, inferring quantitative atmospheric information from  $\Delta^{17}$ O of nitrate requires assessing precisely its controls and to include them into a consistent modeling framework. In the last few years, several models have been proposed to study the spatio-temporal variations of  $\Delta^{17}O$ and relate them to spatio-temporal variations of the fractional contribution of NOx sink reactions. The pioneering work of Lyons (2001) set the stage for the first model study of the seasonal variations of  $\Delta^{17}$ O of atmospheric nitrate by Michalski et al. (2003). Further implementations of  $\Delta^{17}$ O into atmospheric chemistry models were proposed in the following years, from 0-D box-modeling (Morin et al., 2008; Dominguez et al., 2009; Michalski and Xu, 2010) to the 3-D chemical transport model GEOS-Chem (Kunasek et al., 2008; Alexander et al., 2009).

This study revisits some assumptions, hypotheses and approaches previously introduced in the literature (e.g. Michalski et al., 2003; Morin et al., 2007, 2008, 2009; Kunasek et al., 2008; Alexander et al., 2009) and puts them within a consistent framework and perspective that makes it easier to understand and implement in existing atmospheric chemistry models. Limitations of the various assumptions that have been used so far are highlighted and critically evaluated. The overarching goal is to provide a rationale behind assumptions and simplifications that have to be used in large scale model implementation in order to reduce computing costs. The CAABA/MECCA atmospheric chemistry box model (Sander et al., 2011) was used to explicitly calculate the time evolution of the  $\Delta^{17}$ O of short-lived reactive species at each time step. Model runs were performed in a few simple cases to demonstrate the usefulness of such assessments and provide the basis of future analogous studies. Finally, recommendations are given for the implementation of simplifying assumptions into large-scale atmospheric chemistry models.

#### 2 General framework

#### 2.1 The general "mass-balance" equation

The general "mass-balance" equation (also termed the "continuity equation") governing the temporal evolution of the concentration of a given species in a given air parcel is given by:

$$\frac{d}{dt}[\mathbf{X}] = \Sigma_i P_i - \Sigma_j L_j \tag{1}$$

where  $P_i$  and  $L_j$  represent source and sink rates (in cm<sup>-3</sup> s<sup>-1</sup>) of the species X, respectively. Its atmospheric concentration, denoted [X], is expressed in cm<sup>-3</sup>. Sources and sinks include both chemical reactions within the parcel and fluxes at its boundaries. Atmospheric chemistry models are mostly driven by reaction kinetics, so that the chemical components of  $P_i$  and  $L_j$  are simply expressed as a reaction rate constant (usually referred to as *k* values) times the relevant atmospheric concentrations (Jacob, 1999; Finlayson-Pitts and Pitts, 2000).

The implementation of  $\Delta^{17}$ O into the mass balance Eq. (1) follows from mass conservation applied to the oxygen isotope anomaly. Of course, this rather simple method would not apply to isotopic enrichment ( $\delta$ ) values, because isotopic fractionation has to be fully taken into account for every reaction considered (Gromov et al., 2010). The key assumption behind the modeling approach is that sink reactions do not

induce a specific non-mass dependent fractionation, and that every source reaction induces the transfer of a given  $\Delta^{17}$ O value to the newly produced species. Potential contributions of mass-dependent fractionation to the time evolution of  $\Delta^{17}$ O are neglected, because we concentrate on species generally possessing high  $\Delta^{17}$ O values where such effects are negligible (Kaiser et al., 2004). The  $\Delta^{17}$ O mass-balance equation reads:

$$\frac{d}{dt}\left([\mathbf{X}] \times \Delta^{17} \mathbf{O}(\mathbf{X})\right) = \Sigma_i \left(P_i \times \Delta^{17} \mathbf{O}_i(\mathbf{X})\right) - \left(\Sigma_j L_j\right) \times \Delta^{17} \mathbf{O}(\mathbf{X})$$
(2)

where  $\Delta^{17}O(X)$  represents the  $\Delta^{17}O$  of the species X and  $\Delta^{17}O_i(X)$  is the isotope anomaly that is transferred to X through the production channel  $P_i$  of the species X. It is estimated as a function of the  $\Delta^{17}$ O value of the precursors involved in a given production channel for species X, using a mass-balance approach based on the counting of the oxygen atoms transferred throughout a given production channel. Non-mass dependent fractionation induced by a specific reaction can also be taken into account in the equation above. Solving numerically the system of equations formed by Eqs. (1) and (2) for all relevant atmospheric species simultaneously yields the time evolution of the concentration and  $\Delta^{17}$ O of each atmospheric species. This is generally not computationally affordable for large-scale modeling studies such as Alexander et al. (2009). The computation can be carried out for limited periods of time using box models.

#### 2.2 Isotopic exchange reactions

Not only chemical production and destruction impact the  $\Delta^{17}O$  of a given species. Isotopic exchange reactions can also modify it. Their main characteristic is that they have no impact on the chemical budget of a species (i.e., Eq. (1) is not changed), but they have an impact on the isotopic massbalance Eq. (2). The magnitude of an isotopic exchange reaction can be expressed in a similar manner to chemical production or destruction fluxes. In what follows, the rate of the k isotopic exchange reaction is referred to as  $E_k$ ; the ultimate  $\Delta^{17}O$  value that would be attained in species X if the isotopic exchange with the species  $Y_k$  fully proceeds is noted  $\Delta^{17}O(Y_k)$ . Implementing this into Eq. (2) yields:

$$\frac{d}{dt}\left([X] \times \Delta^{17} O(X)\right) = \Sigma_i \left(P_i \times \Delta^{17} O_i(X)\right)$$
(3)  
+ $\Sigma_k \left(E_k \times \Delta^{17} O(Y_k)\right) - \left(\Sigma_j L_j + \Sigma_k E_k\right) \times \Delta^{17} O(X)$ 

#### 2.3 Steady-state approximation

A very commonly used simplification in atmospheric chemistry models is the so-called "photochemical steady-state (PSS)" approximation. This simply assumes that the photolytical lifetime of a given species is sufficiently short that the short-term variations of its concentration are negligible, i.e.  $\frac{d}{dt}[X]\approx 0$ . In other words, a near-perfect balance between sources and sinks for a given species is assumed. Implementing this assumption into the isotopic mass-balance Eq. (2), and taking into account that, at PSS,  $\sum_i P_i = \sum_j L_j$  yields:

$$\Delta^{17} \mathcal{O}(\mathbf{X}) = \frac{\Sigma_i \left( P_i \times \Delta^{17} \mathcal{O}_i(\mathbf{X}) \right)}{\Sigma_i P_i} \tag{4}$$

Stated differently, at PSS the  $\Delta^{17}$ O of a given species is instantaneously equal to the  $\Delta^{17}$ O induced by the combination of its different chemical sources, scaled according to their relative strength, as shown in Eq. (4). Without PSS, the time evolution of the  $\Delta^{17}$ O of a given species must also take into account its  $\Delta^{17}$ O value earlier on. The longer the lifetime of a given species, the slower the time evolution of its  $\Delta^{17}$ O.

### 2.4 Controls on $\Delta^{17}$ O of atmospheric nitrate and hydrogen peroxide H<sub>2</sub>O<sub>2</sub>

#### 2.4.1 Atmospheric nitrate

1. . .

Atmospheric nitrate is formed homogeneously and heterogeneously in the atmosphere through the following reactions (Jacob, 1999; Finlayson-Pitts and Pitts, 2000):

$$NO_2 + OH \xrightarrow{M} HNO_3$$
 (R1)

$$NO_3 + RH \to HNO_3 + R \tag{R2}$$

$$N_2O_5 \xrightarrow{\text{net}} \text{HNO}_3$$
 (R3)

$$HNO_4 \xrightarrow{\text{net}} HNO_3$$
 (R4)

Here, RH represents a generic hydrocarbon. Dry and wet deposition are the main sinks of atmospheric nitrate, controlling its atmospheric lifetime which is on the order of days to weeks (Finlayson-Pitts and Pitts, 2000). A straightforward rearrangement of Eq. (2) yields the equation governing the time evolution of  $\Delta^{17}$ O of atmospheric nitrate:

$$\frac{d}{dt} \left( \left[ \mathrm{NO}_{3}^{-} \right] \times \Delta^{17} \mathrm{O} \left( \mathrm{NO}_{3}^{-} \right) \right)$$

$$= \Sigma_{\mathrm{i}} \left( \mathrm{P}_{\mathrm{i}} \times \Delta^{17} \mathrm{O}_{\mathrm{i}} \left( \mathrm{NO}_{3}^{-} \right) \right) - \frac{\left[ \mathrm{NO}_{3}^{-} \right]}{\tau} \times \Delta^{17} \mathrm{O} \left( \mathrm{NO}_{3}^{-} \right)$$
(5)

where  $\tau$  is the atmospheric lifetime of atmospheric nitrate.  $\Delta^{17}O_i(NO_3^-)$  values can be calculated for each nitrate production channel (Michalski et al., 2003; Morin et al., 2007, 2009; Kunasek et al., 2008):

$$\Delta^{17} \mathcal{O}_{\mathrm{OH}+\mathrm{NO}_2} \left( \mathrm{NO}_3^- \right) \tag{6}$$

$$= 1/3 \times \Delta^{17} \mathrm{O(OH)} + 2/3 \times \Delta^{17} \mathrm{O(NO_2)}$$

$$\Delta^{17} O_{NO_3 + RH} (NO_3^-) = \Delta^{17} O(NO_3)$$
(7)

$$\Delta^{17}O_{N_2O_5 \,\text{hydrol}}(NO_3^-) = 5/6 \times \Delta^{17}O(N_2O_5) \tag{8}$$

$$\Delta^{17} O_{\text{HNO}_4 \text{ hydrol}} \left( \text{NO}_3^- \right) = \Delta^{17} O(\text{HNO}_4) \tag{9}$$

As will be demonstrated below, both the mixing ratio and the  $\Delta^{17}$ O of nitrate precursors vary diurnally. For instance, OH plays a significant role only during the day, and  $\Delta^{17}O(NO_2)$  exhibits a strong diurnal variation with a minimum during the day and a maximum at night. Clearly, only the day-time  $\Delta^{17}O(NO_2)$  values matter for the OH+NO<sub>2</sub> nitrate production channel, since during the night this reaction is suppressed. To account for the fact that the isotopic signature of a given production channel has to be scaled with its strength, we define the diurnally-integrated isotopic signature (DIIS) of the nitrate source, denoted  $\overline{\Delta^{17}O_i(NO_3^-)}$ , as follows:

$$\overline{\Delta^{17}O_{i}(NO_{3}^{-})} = \frac{\int_{0}^{24h} P_{i} \times \Delta^{17}O_{i}(NO_{3}^{-})dt}{\int_{0}^{24h} P_{i}dt}$$
(10)

DIIS values quantify the overall  $\Delta^{17}$ O inherited from a given source reaction, taking into account the scaling of diurnal variations in its strength with the associated  $\Delta^{17}$ O it transfers. Additionally, DIIS is a useful metric to quantify the impact of various environmental settings or hypotheses pertaining to isotopic transfer on the ultimate  $\Delta^{17}$ O of atmospheric nitrate.

In the case where the atmospheric lifetime of a given secondary species is significantly longer than one day, DIIS values can be used to infer the seasonal variations of  $\Delta^{17}$ O from the following equation, virtually assuming that steady-state applies:

$$\Delta^{17} O(NO_3^-) = \frac{\Sigma_i \left( P_i \times \overline{\Delta^{17} O_i \left( NO_3^- \right)} \right)}{\Sigma_i P_i}$$
(11)

Equation (11) takes into account that both  $P_i$  and  $\overline{\Delta^{17}O_i(NO_3^-)}$  values change seasonally or as a function of environmental conditions. This method was implicitly used originally by Michalski et al. (2003) to study the seasonal variations of  $\Delta^{17}O(NO_3^-)$  in coastal California. While correct at the seasonal scale to study seasonal variation of nitrate as far as its lifetime is significantly larger than several days, this method does not adequately address variations of  $\Delta^{17}O(NO_3^-)$  at temporal scale smaller than its atmospheric lifetime (Michalski and Xu, 2010), because sink reactions (both physical and chemical) must then be explicitly taken into account, as shown by Eq. (5).

#### 2.4.2 Hydrogen peroxide (H<sub>2</sub>O<sub>2</sub>)

Hydrogen peroxide is a key atmospheric oxidant which plays a major role for in-cloud oxidation of S(IV) (Finlayson-Pitts and Pitts, 2000; Alexander et al., 2005). Savarino and Thiemens (1999a) demonstrated that it possesses a small but significant  $\Delta^{17}$ O signature, which has then been used to study the partitioning between various S(IV) oxidants in the atmosphere (Savarino et al., 2000; Alexander et al., 2005).  $H_2O_2$  is mostly formed through the self reaction of  $HO_2$ :

$$2HO_2 \rightarrow H_2O_2 + O_2 \tag{R5}$$

The  $\Delta^{17}$ O inherited by H<sub>2</sub>O<sub>2</sub> during the above reaction is equal to  $\Delta^{17}$ O(HO<sub>2</sub>). Indeed, Zhu and Lin (2001) have shown that the most likely mechanism for the self reaction of HO<sub>2</sub> involves a six-member-ring intermediate through headto-tail association, which implies that the two oxygen atoms in H<sub>2</sub>O<sub>2</sub> originate from a single one HO<sub>2</sub> radical out of the two reactants.

The concept of DIIS applies to  $H_2O_2$  in a manner analogous to atmospheric nitrate (see above). Since Reaction (R5) is the sole significant  $H_2O_2$  production pathway, and because the atmospheric lifetime of  $H_2O_2$  is generally larger than one day, the application of Eq. (11) is trivial and shows that seasonal variations of  $\Delta^{17}O(H_2O_2)$  can directly be inferred from variations of  $\overline{\Delta^{17}O_{HO_2+HO_2}(H_2O_2)}$  at first order.

### 3 Material and methods: numerical experiments on $\Delta^{17}O$ of short-lived species

In this study, we focus on the time evolution of the  $\Delta^{17}O$ of short-lived atmospheric reactive species such as  $HO_x$  (=  $OH + HO_2$ ),  $NO_x$  (=  $NO + NO_2$ ) and  $RO_2$ . For simplicity, in this initial study we restrict our analysis to gas-phase reactions and exclude halogen, sulfur and carbonaceous chemistry, to focus on the highly non-linear NO<sub>x</sub>-HO<sub>x</sub>/RO<sub>x</sub>-O<sub>3</sub> chemistry first. The impact of diurnal variations of  $\Delta^{17}O$ of short-lived species on secondary species such as atmospheric nitrate and  $H_2O_2$  is explored through the estimation of diurnally-integrated isotopic signature (DIIS) of these species. This series of 51 reactions (including 14 photolysis reactions) represents a subset of the chemical mechanism implemented in MECCA (Sander et al., 2011) suited for simplified analysis in the remote marine boundary layer. The complete listing of the reactions considered is given as an online supplement to this article. Note that this article mostly seeks to illustrate the basic concepts and equations introduced in Sect. 2, and the impact of performing various hypotheses to propagate the  $\Delta^{17}$ O of ozone throughout atmospheric reactions. This is why heterogeneous reactions leading to the formation of atmospheric nitrate are not explicitly included in the model, because doing so would render the illustration of the main concepts unnecessarily tedious. We concentrate on gas-phase reactions, which provide illustrations of the main behaviors described here, and leave the explicit inclusion of heterogeneous chemistry in the model for a follow-up study comparing observed and simulated diurnal variations of  $\Delta^{17}O(NO_3^-)$ .



**Fig. 1.** Schematic representing the isotopic mass-balance equation including one isotopic exchange reaction. Arrows represent production (or destruction) fluxes. The flux is given by the *P*, *L* and *E* terms, while the corresponding transferred  $\Delta^{17}$ O value is given in brackets. The solid box represent the chemical budget of the species X, while the dashed box takes into account the full isotopic budget of the species X. The scheme illustrates that isotopic exchange reactions have no impact on the chemical budget of a given species X.

#### 3.1 Species with assigned $\Delta^{17}$ O values

#### 3.1.1 $\Delta^{17}O(H_2O)$ and $\Delta^{17}O(O_2)$

Because they represent large oxygen reservoirs with a negligible isotope anomaly, we assume in what follows that  $\Delta^{17}O$  of water vapor (H<sub>2</sub>O) and molecular oxygen (O<sub>2</sub>) are constant, with a value of insignificantly different from 0‰ (Barkan and Luz, 2003, 2005), in comparison to the  $\Delta^{17}O$  of the species dealt with below. Indeed,  $\Delta^{17}O(O_2) = -0.3\%$  (Barkan and Luz, 2003), while  $\Delta^{17}O(H_2O)$  ranges between -1.0 and 0.0‰ (Barkan and Luz, 2005). This simplification is generally made in modeling studies dealing with  $\Delta^{17}O(NO_3^-)$  (Michalski et al., 2003; Alexander et al., 2009), and allows to focus on and interpret the changes in  $\Delta^{17}O$  of reactive and secondary species attributable to chemical transfer of  $\Delta^{17}O$  from ozone.

#### 3.1.2 $\Delta^{17}O(O_3)$

Although several atmospheric reactions induce non-mass dependent fractionation (Brenninkmeijer et al., 2003; Thiemens, 2006) and thus may contribute significantly to the non-zero  $\Delta^{17}$ O values of several atmospheric species, the overwhelming source of non-zero  $\Delta^{17}$ O in the lower atmosphere is ozone (O<sub>3</sub>). As repeatedly mentioned in the recent literature (e.g. Morin et al., 2007; Michalski and Bhattacharya, 2009; Alexander et al., 2009; Dominguez et al., 2009), the tropospheric value of  $\Delta^{17}$ O(O<sub>3</sub>) is controversial. In addition, ozone is isotopically asymmetrical (Janssen, 2005; Marcus, 2008), meaning that the  $\Delta^{17}$ O borne by its terminal and central atoms are different. Furthermore, oxy-

gen atoms of O3 are not chemically equivalent in terms of reactivity during bimolecular reactions. We define  $\Delta^{17}O(O_2^{\star})$ as the  $\Delta^{17}$ O value which is transferred along with the terminal O atom of ozone. The motivation for this choice is two-fold. Firstly, terminal oxygen atoms generally have a greater probability of being transferred than the central oxygen atom (Savarino et al., 2008), and this probability even equals 1 for the following reactions: O<sub>3</sub>+Ag<sub>metal</sub> (Bhattacharya et al., 2008),  $O_3 + NO_2^-$  (Liu et al., 2001; Michalski and Bhattacharya, 2009), O<sub>3</sub>+NO<sub>2</sub> (Peiro-Garcia and Nebot-Gil, 2003, Berhanu and Bhattacharya, personal communication, 2011). Secondly, we note that atmospheric direct measurements of  $\Delta^{17}O(O_2^{\star})$  are now possible using the reaction O<sub>3</sub>+NO<sub>2</sub><sup>-</sup> as a chemical probe (Michalski and Bhattacharya, 2009; Vicars et al., 2011).  $\Delta^{17}O(O_3^*)$ , which may be referred to as the "transferrable  $\Delta^{17}O(O_3)$  values through bimolecular reactions operating through the terminal O atom of ozone", can now by directly measured.

The link between  $\Delta^{17}O(O_3)$  and  $\Delta^{17}O(O_3^*)$  is complex due to our currently limited knowledge of the intramolecular distribution of  $\Delta^{17}O$  within the ozone molecule. For  $\Delta^{17}O(O_3)$  values ranging between 20 and 40‰, currently existing experimental data yield the following relationship:

$$\Delta^{17} \mathcal{O}(\mathcal{O}_3^{\star}) = 1.5 \times \Delta^{17} \mathcal{O}(\mathcal{O}_3) \tag{12}$$

We note that, although this is not entirely consistent with the full range of experimental observations and their theoretical implications, this equation is equivalent to considering that  $\Delta^{17}$ O resides exclusively on the terminal oxygen atoms of ozone (Michalski and Bhattacharya, 2009).

Consistent with the modeling studies carried out hitherto (Michalski et al., 2003; Morin et al., 2008; Kunasek et al., 2008; Alexander et al., 2009; Dominguez et al., 2009; Michalski and Xu, 2010), we use a constant value of  $\Delta^{17}O(O_3)$  of 30%, which falls within the 25–35% range generally found in the literature (see Brenninkmeijer et al. (2003) or Morin et al. (2007) for details), to illustrate the results of the calculations. Most results of the present work can be extrapolated to different  $\Delta^{17}O(O_3)$  values simply by scaling them with the ratio of a different value of  $\Delta^{17}O(O_3)$ to our current choice. Note that while we have chosen a constant value for  $\Delta^{17}O(O_3)$  to remain consistent with previous studies, the model can account for temporal variations of  $\Delta^{17}O(O_3)$  (Vicars et al., 2011). Should  $\Delta^{17}O(O_3)$  exhibit significant diurnal variations, the quantitative results provided below for illustration purposes would need to be revised, but the underlying concepts exposed here would remain valid. In this case, only an explicit modeling framework as described here would then be suitable to compute the  $\Delta^{17}$ O values of reactive and secondary atmospheric species.

In this work, bimolecular reactions involving ozone are:

$$NO + O_3 \rightarrow NO_2 + O_2 \tag{R6}$$

$$NO_2 + O_3 \rightarrow NO_3 + O_2 \tag{R7}$$

| Rxn #          |   | $\Delta^{17}O_i(X)$  |
|----------------|---|--|
| G3109          | $NO_3 + NO_2 \rightarrow N_2O_5$  | $\Delta^{17}O_{G3109}(N_2O_5 - NO_2) = \Delta^{17}O(NO_2)$<br>$\Delta^{17}O_{G3109}(N_2O_5 - NO_3) = \Delta^{17}O(NO_3)$ |
| G3200          | NO+OH→HONO  | $\Delta^{17}O_{G3200}(HONO) = 1/2 \left( \Delta^{17}O(NO) + \Delta^{17}O(OH) \right)$                                    |
| G3202          | $NO_2+OH \rightarrow HNO_3$   | $\Delta^{17}O_{G3202}(HNO_3) = 1/3 (2\Delta^{17}O(NO_2) + \Delta^{17}O(OH))$   |
| G3203<br>G4109 | $\begin{array}{l} \text{NO}_2 + \text{HO}_2(+\text{M}) \rightarrow \text{HNO}_4 \\ \text{HCHO} + \text{NO}_3(+\text{O}_2) \rightarrow \text{HNO}_3 + \text{CO} + \text{HO}_2 \end{array}$ | $\Delta^{17}O_{G3203}(HNO_4 - NO_2) = \Delta^{17}O(NO_2)$<br>$\Delta^{17}O_{G4109}(HNO_3) = \Delta^{17}O(NO_3)$          |

**Table 1.** Listing of  $\Delta^{17}O_i(X)$  relevant to Case 1 (PSS NO<sub>x</sub>), if not given in Sect. 3.2.1. Note the special case of HNO<sub>4</sub> and N<sub>2</sub>O<sub>5</sub>, for which the  $\Delta^{17}O$  of each of the molecule making up the dimer are explicitly referred to and tracked.

$$OH + O_3 \rightarrow HO_2 + O_2 \tag{R8}$$

$$\mathrm{HO}_2 + \mathrm{O}_3 \to \mathrm{OH} + 2\mathrm{O}_2 \tag{R9}$$

Ab initio calculations on the mechanism of the reaction  $NO_2+O_3$  reveal that it proceeds through the abstraction of the terminal O atom of ozone (Peiro-Garcia and Nebot-Gil, 2003). In addition, the transfer of  $\Delta^{17}O$  through reaction  $NO_2+O_3$  has been recently studied in the lab and it was also concluded that only the terminal O atom of ozone is transferred (Berhanu and Bhattacharya, personal communication, 2011). We also assume that the OH+O<sub>3</sub> reaction proceeds exclusively through the transfer of the terminal O atom of ozone. We thus apply  $\Delta^{17}O(O_3^*)$  as the isotopic signature of these reactions. Considering a  $\Delta^{17}O(O_3)$  value of 30‰, this corresponds to  $\Delta^{17}O(O_3^*) = 45\%$ . In the case of the reaction NO+O<sub>3</sub>, we apply the  $\Delta^{17}O$  transfer rate determined experimentally by Savarino et al. (2008):

$$\Delta^{17}O_{\rm NO+O_3}(\rm NO_2) = 1.18 \times \Delta^{17}O(\rm O_3) + 6.6 \times 10^{-3}$$
(13)

With a  $\Delta^{17}O(O_3)$  value of 30‰, this corresponds to a  $\Delta^{17}O_{NO+O_3}(NO_2)$  value of 42‰. The fact that the  $\Delta^{17}O(O_3)$  transferred through this reaction is lower than through reactions described before stems from the fact that not only the terminal, but also the central oxygen atom of ozone reacts with NO, as described in detail by Savarino et al. (2008). Lastly, in the case of reaction HO<sub>2</sub>+O<sub>3</sub> this is irrelevant to the value of  $\Delta^{17}O(OH)$  because the oxygen atom in OH stems from HO<sub>2</sub>, not ozone.

#### $3.1.3 \quad \Delta^{17}O(RO_2)$

In this study we explicitly separate HO<sub>2</sub> from other peroxy radicals, denoted RO<sub>2</sub> where R represents a carbonaceous chain, because the chemical budget of HO<sub>2</sub> and RO<sub>2</sub> is very different. While reactions involving ozone contribute to the budget of HO<sub>2</sub>, the only source of RO<sub>2</sub> is the reaction between O<sub>2</sub> and a R radical. Since  $\Delta^{17}O(O_2)=0\%$  (see above), this immediately implies that  $\Delta^{17}O(RO_2)=0\%$  under all tropospheric conditions.

### 3.2 Overview of the sets of hypotheses regarding the $\Delta^{17}$ O transfer throughout chemical reactions

Below we present the various sets of hypotheses (numbered Cases 1 to 6) implemented to compute the time evolution of  $\Delta^{17}O$  of the species of interest using various assumptions in terms of  $\Delta^{17}O$  transfer.

### 3.2.1 Case 1: NO<sub>x</sub> photochemical steady-state (PSS) and basic hypotheses

- OH, HO<sub>2</sub>:  $\Delta^{17}$ O of both species is equal to 0‰.
- NO, NO<sub>2</sub>, NO<sub>3</sub>: The  $\Delta^{17}$ O of these species is calculated using PSS:

$$\Delta^{17}O(NO_2) = \alpha \times \Delta^{17}O_{NO+O_3}(NO_2)$$

with

$$\alpha = \frac{k_{\text{NO+O_3}}[\text{O_3}]}{k_{\text{NO+O_3}}[\text{O_3}] + k_{\text{NO+HO_2}}[\text{HO_2}] + k_{\text{NO+RO_2}}[\text{RO_2}]}$$

as defined by Michalski et al. (2003) and demonstrated in Morin et al. (2007). It follows from PSS that  $\Delta^{17}O(NO) = \Delta^{17}O(NO_2)$ .  $\Delta^{17}O(NO_3)$  is given as:

$$\Delta^{17}O(NO_3) = 1/3 \left( 2\alpha \Delta^{17}O_{NO+O_3}(NO_2) + \Delta^{17}O(O_3^{\star}) \right)$$

- HONO, HNO<sub>3</sub>, H<sub>2</sub>O<sub>2</sub>: the calculation of  $\Delta^{17}$ O is calculated following Eq. (2).
- N<sub>2</sub>O<sub>5</sub>, HNO<sub>4</sub>: both these species are dimers formed by the combinations of two radicals (NO<sub>2</sub> and NO<sub>3</sub>, and NO<sub>2</sub> and HO<sub>2</sub>, respectively). This makes necessary to track the time evolution of the  $\Delta^{17}$ O of both components of the dimer making up N<sub>2</sub>O<sub>5</sub> and HNO<sub>4</sub>, respectively, since their O atoms are not isotopically equivalents.

Table 1 gives the  $\Delta^{17}O_i(X)$  values for each species produced, if different from 0 and not given above.

| Rxn #  |  | $\Delta^{17}O_i(X)$   |
|--------|--|---|
| G3103  | $NO+O_3 \rightarrow NO_2+O_2$                      | $\Delta^{17}O_{G3103}(NO_2) = 1/2 \left( \Delta^{17}O_{NO+O_3}(NO_2) + \Delta^{17}O(NO) \right)$              |
| G3106  | $NO_2+O_3 \rightarrow NO_3+O_2$                    | $\Delta^{17}O_{G3106}(NO_3) = 1/2 \left( \Delta^{17}O(O_3^{\star}) + \Delta^{17}O(NO_2) \right)$              |
| G3108  | $NO_3 + NO \rightarrow 2NO_2$                      | $\Delta^{17}O_{G3108}(NO_2) = 1/2 \left( \Delta^{17}O(NO_3) + \Delta^{17}O(NO) \right)$                       |
| G3110  | $N_2O_5(+M) \rightarrow NO_2 + NO_3$               | $\Delta^{17}O_{G3110}(NO_2) = 1/2 \left( \Delta^{17}O(N_2O_5 - NO_2) + \Delta^{17}O(N_2O_5 - NO_3) \right)$   |
|        |  | $\Delta^{17}O_{G3110}(NO_3) = 1/3 \left( \Delta^{17}O(N_2O_5 - NO_2) + 2\Delta^{17}O(N_2O_5 - NO_3) \right)$  |
| G3201  | $NO+HO_2 \rightarrow NO_2+OH$                      | $\Delta^{17}O_{G3201}(NO_2) = 1/2 \left( \Delta^{17}O(HO_2) + \Delta^{17}O(NO) \right)$                       |
| G3204  | $NO_3+HO_2\rightarrow NO_2+OH+O_2$                 | $\Delta^{17}O_{G3204}(NO_2) = \Delta^{17}O(NO_3)$   |
| G3205  | $HONO+OH \rightarrow NO_2+H_2O$                    | $\Delta^{17}O_{G3205}(NO_2) = \Delta^{17}O(HONO)$   |
| G3206  | $HNO_3+OH \rightarrow H_2O+NO_3$                   | $\Delta^{17}O_{G3206}(NO_3) = \Delta^{17}O(HNO_3)$  |
| G3207  | $HNO_4(+M) \rightarrow NO_2 + HO_2$                | $\Delta^{17}O_{G3207}(NO_2) = \Delta^{17}O(HNO_4 - NO_2)$   |
| G3208  | $HNO_4+OH \rightarrow NO_2+H_2O+O_2$               | $\Delta^{17}O_{G3208}(NO_2) = \Delta^{17}O(HNO_4 - NO_2)$   |
| G4104  | $CH_3O_2 + NO \rightarrow HCHO + NO_2 + HO_2$      | $\Delta^{17}O_{G4104}(NO_2) = 1/2\Delta^{17}O(NO)$  |
| G4105  | $CH_{3}O_{2}+NO_{3}\rightarrow HCHO+HO_{2}+NO_{2}$ | $\Delta^{17}O_{G4105}(NO_2) = \Delta^{17}O(NO_3)$   |
| J3101  | $NO_2 + h\nu \rightarrow NO + O(^3P)$              | $\Delta^{17}O_{J3101}(NO) = \Delta^{17}O(NO_2)$   |
| J3103a | $NO_3 + h\nu \rightarrow NO_2 + O(^3P)$            | $\Delta^{17}O_{J3103a}(NO_2) = \Delta^{17}O(NO_3)$  |
| J3103b | $NO_3 + h\nu \rightarrow NO + O_2$                 | $\Delta^{17}O_{J3103b}(NO) = \Delta^{17}O(NO_3)$  |
| J3104a | $N_2O_5+h\nu \rightarrow NO_2+NO_3$                | $\Delta^{17}O_{J3104a}(NO_2) = 1/2 \left( \Delta^{17}O(N_2O_5 - NO_2) + \Delta^{17}O(N_2O_5 - NO_3) \right)$  |
|        |  | $\Delta^{17}O_{J3104a}(NO_3) = 1/3 \left( \Delta^{17}O(N_2O_5 - NO_2) + 2\Delta^{17}O(N_2O_5 - NO_3) \right)$ |
| J3200  | $HONO+h\nu \rightarrow OH+NO$                      | $\Delta^{17}O_{J3200}(NO) = \Delta^{17}O(HONO)$   |
| J3201  | $HNO_3 + h\nu \rightarrow OH + NO_2$               | $\Delta^{17}O_{J3201}(NO_2) = \Delta^{17}O(HNO_3)$  |
| J3202  | $HNO_4 + h\nu \rightarrow 0.667NO_2 + 0.667HO_2$   | $\Delta^{17}O_{J3202}(NO_2) = \Delta^{17}O(HNO_4 - NO_2)$   |
|        | +0.333NO <sub>3</sub> +0.333OH                     | $\Delta^{17}O_{J3202}(NO_3) = 2/3 \Delta^{17}O(HNO_4 - NO_2) + 1/3 \Delta^{17}O(HNO_4 - HO_2)$                |

**Table 2.**  $\Delta^{17}O_i(X)$  for Case 2 (explicit NO<sub>x</sub>). Only equations featuring different  $\Delta^{17}O_i(X)$  than in Case 1 (PSS NO<sub>x</sub>) are presented here.

#### 3.2.2 Case 2: explicit NO<sub>x</sub>

In this case, the time evolution of  $\Delta^{17}$ O of NO, NO<sub>2</sub> and NO<sub>3</sub> is computed explicitly. This means that the PSS approximation is not used at all regarding the computation of the  $\Delta^{17}$ O of these species. Table 2 presents the chemical reactions for which the  $\Delta^{17}$ O<sub>i</sub>(X) is different than in Case 1 (PSS NO<sub>x</sub>).

The decomposition of N<sub>2</sub>O<sub>5</sub> and HNO<sub>4</sub> through photolysis or thermal decomposition is treated as follows. Our base hypothesis is that the photolysis and thermal decomposition of HNO<sub>4</sub> proceed without scrambling of its isotopic composition upon dissociation. The case of N<sub>2</sub>O<sub>5</sub> is more complex and is illustrated by Fig. 2. Upon dissociation or photolysis of N<sub>2</sub>O<sub>5</sub>, the  $\Delta^{17}$ O of NO<sub>2</sub> and NO<sub>3</sub> formed is given by the following equations, based on the assumption that the two N–O bonds of the dimer have an equal probability to break:

$$\begin{split} &\Delta^{17} O_{N_2O_5 decomp}(NO_2) \\ &= 1/2 \times \left( \Delta^{17} O(N_2O_5 - NO_2) + \Delta^{17} O(N_2O_5 - NO_3) \right) \\ &\Delta^{17} O_{N_2O_5 decomp}(NO_3) \\ &= 1/3 \times \left( \Delta^{17} O(N_2O_5 - NO_2) + 2 \times \Delta^{17} O(N_2O_5 - NO_3) \right) \end{split}$$

where  $\Delta^{17}O(N_2O_5 - NO_2)$  and  $\Delta^{17}O(N_2O_5 - NO_3)$  represent the  $\Delta^{17}O$  of the two dimers making up N<sub>2</sub>O<sub>5</sub>, i.e. NO<sub>2</sub> and NO<sub>3</sub>, respectively.

#### 3.2.3 Case 3: explicit HO<sub>x</sub>

In this case, in addition to the hypotheses of Case 2 (explicit NO<sub>x</sub>), the  $\Delta^{17}$ O value of HO<sub>2</sub> is allowed to vary in time and is computed explicitly. This also induces non-zero values of  $\Delta^{17}O(H_2O_2)$ , through reaction HO<sub>2</sub>+HO<sub>2</sub>. Table 3 presents the chemical reactions for which the  $\Delta^{17}O_i(X)$  is different than in Case 1 (PSS  $NO_x$ ) and 2 (explicit  $NO_x$ ). Note that in Case 3  $\Delta^{17}O(OH)$  is assigned a value of to 0‰. Note that, strictly speaking, the O atoms in HO<sub>2</sub> are chemically not equivalent, because one is bonded to the H atom while the other one is not. This opens the possibility that the intramolecular distribution of  $\Delta^{17}$ O within HO<sub>2</sub> is not statistically distributed, in a manner somewhat equivalent to ozone, as first noted by Savarino and Thiemens (1999b). For the sake of simplicity, we consider here that the  $\Delta^{17}O$  of HO<sub>2</sub> can be represented as a single value, implicitly assuming that the intramolecular distribution of HO<sub>2</sub> is statistical.



**Fig. 2.** Schematic illustrating the fate of oxygen atoms from NO<sub>3</sub> (in red) and NO<sub>2</sub> (in blue) during the formation (a) and decomposition (b) of N<sub>2</sub>O<sub>5</sub>. Owing to the symmetry of the N<sub>2</sub>O<sub>5</sub> molecule, the N–O bonds are expected to have an equal probability to break upon thermal decomposition or photolysis. The  $\Delta^{17}O$  of NO<sub>2</sub> formed upon this dissociation corresponds to the average between the  $\Delta^{17}O$  of NO<sub>2</sub> and NO<sub>3</sub> making up N<sub>2</sub>O<sub>5</sub>. The  $\Delta^{17}O$  of NO<sub>3</sub> formed upon the dissociation of N<sub>2</sub>O<sub>5</sub> is a slightly more complex function, due to the fact that in half cases two out of the three oxygen atoms of NO<sub>3</sub> initially stem from NO<sub>2</sub>, while one comes from the original NO<sub>3</sub>. See Sect. 3.2.2 for details.

#### 3.2.4 Additional tests

In addition to the three main cases presented above, three additional tests were performed. They all are based on Case 3 (explicit  $NO_x$  and  $HO_x$ ), i.e. they can all be independently compared to Case 3.

#### Case 4: NMDF H+O<sub>2</sub>

In this case, we take into account that reaction  $H+O_2 \rightarrow HO_2$ induces non-mass dependent fractionation of oxygen isotopes. The effect is assumed to be on the order of 1‰, according to Savarino and Thiemens (1999b). In practice, any  $HO_2$  produced through this channel is thus attributed a  $\Delta^{17}O$ value of 1‰.

#### Case 5: Scrambling upon the decomposition of HNO<sub>4</sub>

In this case, it is assumed that the thermal decomposition and photolysis of HNO<sub>4</sub> induce a scrambling of its oxygen atoms. Table 4 presents the chemical reactions for which the  $\Delta^{17}O_i(X)$  is different than in Case 1 (PSS NO<sub>x</sub>) for the species produced upon the thermal decomposition of HNO<sub>4</sub>.

#### Case 6: isotopic exchange between OH and H<sub>2</sub>O

In this case, we take into account the isotopic exchange reaction between OH and  $H_2O$ :

$$QH + H_2^{16}O \to {}^{16}OH + H_2Q$$
 (R10)

where Q denotes one of the three O isotopes. This reactions leads to the erasion of  $\Delta^{17}O(OH)$  following isotopic exchange with water vapor, which has a 0‰  $\Delta^{17}O$ . Tropospheric OH is always at photochemical steady-state during daytime given its extremely short lifetime (a few seconds at most). Under such conditions, its  $\Delta^{17}O$  is computed as follows:

$$\Delta^{17} O(OH) = \frac{\Sigma_i L_i}{\Sigma_i L_i + k_{R10} [OH] [H_2 O]} \times \Delta^{17} O_{OHsource} (OH)$$
(14)

where values for  $k_{\rm R10}$  were measured by Dubey et al. (1997). In this study, the sole chemical reaction considered inducing non-zero  $\Delta^{17}$ O values in OH is the reaction between O(<sup>1</sup>D) and H<sub>2</sub>O. Mass-balance states that  $\Delta^{17}O_{\rm OHsource}(\rm OH) =$  $1/2 \times \Delta^{17}O(O_3^{\star})$  (Morin et al., 2007). Under most conditions prevailing in the lower troposphere in mid-latitudes,  $\Delta^{17}O(\rm OH) = 0\%$  (Michalski et al., 2003). However, under cold conditions the isotopic exchange reaction can compete with its OH chemical sinks (Morin et al., 2007), which are mostly OH+CH<sub>4</sub> and OH+CO (Finlayson-Pitts and Pitts, 2000). When Case 6 is tested,  $\Delta^{17}O(\rm OH)$  is assigned a value calculated from Eq. (14) and the mixing ratio of CO, CH<sub>4</sub> and H<sub>2</sub>O and the relevant kinetic rate constants.

### 3.3 Numerical implementation and computation of $\Delta^{17}O$

#### MECCA

The chemistry module MECCA (Model Efficiently Computing the Chemistry of the Atmosphere) is embedded in the CAABA (Chemistry As A Box-model Application) boxmodel (Sander et al., 2011). It uses an adaptative time resolution mathematical method to solve the stiff set of equations describing the evolution of the chemical composition of the portion of atmosphere hypothetically contained in a closed box.

#### **Isotopic equations**

The MECCA chemistry module, like all atmospheric chemistry models, solves the continuity equation for all considered species and all reactions simultaneously:

$$\frac{d}{dt}[\mathbf{X}] = \Sigma_i P_i - \Sigma_j L_j$$

Considering that the model provides the necessary data at a time step *t*, it follows that at the next time step  $t+\Delta t$ :

$$[X](t + \Delta t) = [X](t) + \Delta t \times \Sigma_i P_i - \Delta t \times \Sigma_j L_j$$
(15)

| Rxn # |   | $\Delta^{17}O_i(X)$   |
|-------|---|---|
| G2104 | $OH+O_3 \rightarrow HO_2+O_2$   | $\Delta^{17}O_{G2104}(HO_2) = 1/2 \times \Delta^{17}O(O_3^{\star})$ |
| G2110 | $2HO_2 \rightarrow H_2O_2 + O_2$  | $\Delta^{17}O_{G2110}(H_2O_2) = \Delta^{17}O(HO_2)$                 |
| G2112 | $H_2O_2 + OH \rightarrow H_2O + HO_2$   | $\Delta^{17}O_{G2112}(HO_2) = \Delta^{17}O(H_2O_2)$                 |
| G3203 | $NO_2 + HO_2(+M) \rightarrow HNO_4$   | $\Delta^{17}O_{G3203}(HNO_4 - HO_2) = \Delta^{17}O(HO_2)$           |
| G3207 | $HNO_4(+M) \rightarrow NO_2 + HO_2$   | $\Delta^{17}O_{G3207}(HO_2) = \Delta^{17}O(HNO_4 - HO_2)$           |
| J3202 | $\begin{array}{c} HNO_4 + h\nu \! \rightarrow \! 0.667 NO_2 \! + \! 0.667 HO_2 \\ + \! 0.333 NO_3 \! + \! 0.333 OH \end{array}$ | $\Delta^{17}O_{J3202}(HO_2) = \Delta^{17}O(HNO_4 - HO_2)$           |

**Table 3.**  $\Delta^{17}O_i(X)$  for Case 3: explicit NO<sub>x</sub> and HO<sub>x</sub>. Only equations featuring  $\Delta^{17}O_i(HO_2)$  different from 0‰ are presented here.

**Table 4.**  $\Delta^{17}O_i(X)$  for Case 5: Scrambling upon thermal decomposition and photolysis of HNO<sub>4</sub>

| Rxn # |  | $\Delta^{17}O_i(X)$   |
|-------|--|---|
| G3207 | $HNO_4(+M) \rightarrow NO_2 + HO_2$              | $\Delta^{17}O_{G3207}(NO_2) = 1/2 \left( \Delta^{17}O(HNO_4 - NO_2) + \Delta^{17}O(HNO_4 - HO_2) \right)$ |
|       |  | $\Delta^{17}O_{G3207}(HO_2) = 1/2 \left( \Delta^{17}O(HNO_4 - NO_2) + \Delta^{17}O(HNO_4 - HO_2) \right)$ |
| J3202 | $HNO_4 + h\nu \rightarrow 0.667NO_2 + 0.667HO_2$ | $\Delta^{17}O_{J3202}(NO_2) = 1/2 \left( \Delta^{17}O(HNO_4 - NO_2) + \Delta^{17}O(HNO_4 - HO_2) \right)$ |
|       | +0.333NO <sub>3</sub> +0.333OH                   | $\Delta^{17}O_{J3202}(NO_3) = 1/2 \left( \Delta^{17}O(HNO_4 - NO_2) + \Delta^{17}O(HNO_4 - HO_2) \right)$ |
|       |  | $\Delta^{17}O_{J3202}(HO_2) = 1/2 \left( \Delta^{17}O(HNO_4 - NO_2) + \Delta^{17}O(HNO_4 - HO_2) \right)$ |

The same applies to the isotopic continuity equation (Eq. 2), so that:

 $[\mathbf{X}](t + \Delta t) \times \Delta^{17} \mathbf{O}(\mathbf{X})(t + \Delta t) = [\mathbf{X}](t) \times \Delta^{17} \mathbf{O}(\mathbf{X})(t)$ (16)

$$+\Delta t \times \Sigma \left( P_i \times \Delta^{17} \mathcal{O}_i(\mathbf{X})(t) \right) - \Delta t \times \left( \Sigma_j L_j \right) \times \Delta^{17} \mathcal{O}(\mathbf{X})(t)$$

By combining Eqs. (15) and (3.3),  $\Delta^{17}O(X)(t+\Delta t)$  can be inferred as a function of the relevant chemical and isotopic data at time *t*. This simple explicit approach was implemented in a computer program separate from MECCA, which takes as input a data file containing the variables dealt with in Eqs. (15) and (3.3) at each time step, and processes them according to the different cases described above in terms of  $\Delta^{17}O_i(X)$  to compute the time evolution of  $\Delta^{17}O$ of each relevant species.

One major issue that has to be considered when using this simple approach pertains to the comparison of the chemical lifetime of a species X and the time-step of the integration of Eqs. (15) and (3.3). Indeed, if the lifetime of X is shorter than the time step considered, then the total chemical production or destruction during a given time step  $\Delta t$  may exceed the amount of species X dealt with in the box (or grid-cell). This causes immediate failure of the integration procedure. The time step of the isotopic calculations performed here was chosen accordingly. A time resolution of 10 s was found to be sufficient to avoid integration issues such as described above. More integrated approaches, fully embedded into the box-model itself, have been developed and avoid such shortcomings (see e.g., Gromov et al., 2010). However, for the sake of

the present study, and taking advantage of the easiness of manipulating  $\Delta^{17}$ O through simple mass-balance equations, we preferred the implementation presented above for this study.

#### 3.4 Presentation of MECCA model runs

Our base-run corresponds to atmospheric settings typical of the remote, mid-latitude (45° N) boundary layer during springtime. Photolysis rate coefficients are calculated by the model (Sander et al., 2005). The model run is started on 1 April, at a temperature of 293 K, a relative humidity of 81%, with a starting NO<sub>2</sub> mixing ratio of  $20 \,\mathrm{pmol}\,\mathrm{mol}^{-1}$ . Initial values for the mixing-ratio of main atmospheric species follow: CH<sub>4</sub>,  $1.8 \,\mu$ mol mol<sup>-1</sup>; CO, 70 nmol mol<sup>-1</sup>;  $H_2O_2$ , 600 pmol mol<sup>-1</sup>; HNO<sub>3</sub>, 5 pmol mol<sup>-1</sup>; HCHO,  $30 \text{ pmol mol}^{-1}$ , O<sub>3</sub>, 25 nmol mol<sup>-1</sup>. The model accounts for dry deposition of NO<sub>2</sub>, HNO<sub>3</sub>, N<sub>2</sub>O<sub>5</sub> and H<sub>2</sub>O<sub>2</sub>. In contrast, no emissions into the model box are considered during the model run. After a spin-up time of 1 day, sufficient to initialize the mixing ratio of short-lived species, the time evolution of the mixing ratio and isotope anomaly of short-lived species is analyzed during 36h, corresponding to the time frame between 24 and 60 h from the start of the model run. In lack of emissions of primary species into the box considered, this suffices to identify and study the main features of the diurnal variations of the mixing ratio and  $\Delta^{17}$ O of shortlived species while not suffering from the inherent limitations of box-modeling experienced when longer time periods are considered (Sander et al., 2005).



**Fig. 3.** Time series of the mixing ratio  $(\chi)$  of the main atmospheric species studied, for the base model run (T=293 K, 45° N, see Sect. 3.4 for details). (a) OH and HO<sub>2</sub>, (b) O<sub>3</sub> and H<sub>2</sub>O<sub>2</sub>, (c) N<sub>2</sub>O<sub>5</sub>, HNO<sub>4</sub>, HNO<sub>3</sub> and HONO, (d) NO, NO<sub>2</sub> and NO<sub>3</sub>. The shaded area represents the time period used in the rest of article for the isotopic discussion.

We concentrate our in-depth analysis on this one model run, which shows many different interesting features of the diurnal variations of  $\Delta^{17}$ O of short-lived species and their sensitivity of the various assumptions tested through Cases 1 to 6. The intricacy and the highly non-linear coupling between HO<sub>x</sub> and NO<sub>x</sub> including their  $\Delta^{17}$ O requires careful attention to decipher the causes for the variation of  $\Delta^{17}$ O. The relevance of the conclusions reached from this analysis is assessed using other model runs undertaken under different atmospheric conditions. Indeed, atmospheric chemical processes depend in particular on temperature, time of the year and latitude (through their control of incoming solar radiation) and the chemical regime of the atmosphere.

#### 4 Results

#### 4.1 Description and analysis of the base model run

### 4.1.1 Diurnal variations in mixing ratios and reaction rates

Figure 3 shows the evolution of the mixing ratio of O<sub>3</sub>, HO<sub>x</sub> and  $NO_x/NO_y$  (NO<sub>y</sub> refers to sum of NO<sub>x</sub> and its reservoir species: NO<sub>3</sub>, N<sub>2</sub>O<sub>5</sub>, HONO, HNO<sub>3</sub>, HNO<sub>4</sub>) simulated by MECCA under the conditions of the base model run presented in Sect. 3.4. Negative time corresponds to the abovementioned spin-up time period. Note that the time axis is the same for all the plots exhibiting time series in this article. The model results show typical variations in the mixing ratio of the species of interest, notably with peak values of OH, HO<sub>2</sub> and NO reached during the day. The  $NO_x/NO_y$ partitioning changes diurnally, with species such as NO<sub>3</sub> and N<sub>2</sub>O<sub>5</sub> present mostly during the night, and species such as HNO<sub>4</sub>, HONO present mostly during the day. H<sub>2</sub>O<sub>2</sub> is produced during the day, and undergoes dry deposition which leads to a reduction of its mixing ratio during the night. The mixing ratio of ozone remains quasi-constant during the time period studied, illustrating that the simulation reproduces the chemical steady-state prevailing in the remote, mid-latitude boundary layer.

As a straightforward corollary of the above paragraph, it appears that in this simulation the OH+NO<sub>2</sub> and HNO<sub>4</sub> hydrolysis are mostly daytime nitrate production pathways, while NO<sub>3</sub>+RH and N<sub>2</sub>O<sub>5</sub> hydrolysis proceed only at night, when significant amounts of NO<sub>3</sub> and N<sub>2</sub>O<sub>5</sub> are present. Note that this study does not aim at disentangling complex aspects of the daytime chemistry of N<sub>2</sub>O<sub>5</sub> revealed by recent field campaigns (e.g., Brown et al., 2006). H<sub>2</sub>O<sub>2</sub> is only produced during the day, when HO<sub>2</sub> maximizes. Figure 4 exemplifies such opposed behavior and illustrates the concept behind diurnally-integrated isotopic signature (DIIS) of the nitrate and hydrogen peroxide sources. From the analysis of this figure, it appears obvious that the nighttime  $\Delta^{17}$ O values of HO<sub>2</sub> have no impact on the  $\Delta^{17}$ O of H<sub>2</sub>O<sub>2</sub> produced. Only daytime  $\Delta^{17}$ O values are worth discussing in this case.

## 4.1.2 Overview of the diurnally-integrated isotopic signatures (DIIS) values for atmospheric nitrate and hydrogen peroxide sources

Table 5 shows the DIIS  $(\Delta^{17}O_i(NO_3^-))$  values for the four atmospheric nitrate sources considered, as well as for H<sub>2</sub>O<sub>2</sub>. The main results of this simulation are that

 $<sup>- \</sup>Delta^{17}O_{OH+NO_2}(NO_3^-)$  values are on the order of 20‰ and do not seem significantly dependent to within 0.1‰ upon the different isotopic assumptions tested, except Case 5 (Isotopic scrambling during thermal decomposition and photolysis of HNO<sub>4</sub>) which is detailed below. In particular, whether  $\Delta^{17}O(NO_2)$  is computed

**Table 5.** Diurnally-integrated isotopic signature (DIIS) values for the atmospheric nitrate and  $H_2O_2$  production channels, respectively (in % units). Overview of the results from the base model run carried out in springtime (1 April) under mid-latitude (45° N) remote boundary layer conditions.

|   | $OH + NO_2$ | $NO_3 + RH$ | N <sub>2</sub> O <sub>5</sub> hydrol | HNO4 hydrol | $H_2O_2$ |
|---|-------------|-------------|--------------------------------------|-------------|----------|
| Case $1 - NO_x PSS$                     | 20.3        | 42.6        | 35.1                                 | 15.7        | 0.0      |
| Case $2 - NO_x$ explicit                | 20.4        | 40.2        | 33.2                                 | 15.7        | 0.0      |
| Case $3 - HO_x$ explicit                | 20.5        | 40.2        | 33.3                                 | 16.3        | 1.1      |
| Case 4 – MIF in $H + O_2$               | 20.5        | 40.2        | 33.3                                 | 16.7        | 1.8      |
| Case 5 – Scrambling of HNO <sub>4</sub> | 19.7        | 38.0        | 31.2                                 | 16.0        | 1.5      |
| Case $6 - OH + H_2O$ isotopic exhange   | 20.5        | 40.2        | 33.3                                 | 16.3        | 1.1      |
|   |             |             |                                      |             |          |

explicitly or using the PSS formula has no impact on the DIIS value. This shows that, to compute the DIIS of this reaction channel, it is correct to use the PSS formalism to calculate  $\Delta^{17}O(NO_2)$ . This stems from the fact that, when  $\Delta^{17}O(NO_2)$  deviates from its PSS value, i.e. during the night, the OH+NO\_2 reaction is ineffective owing to the simultaneous absence of OH. Thus, PSS  $\Delta^{17}O(NO_2)$  values can be used to compute the DIIS of this reaction channel, as long as they are correctly scaled with the strength of the OH+NO\_2 reaction.

- $\Delta^{17}O_{HNO_4 hydrol} (NO_3^-) \text{ values are on the order of } 16\% \\ \text{ and do not depend on the method to compute the } \Delta^{17}O \\ \text{ of } NO_x \text{ (Case 1 (PSS NO_x) and Case 2 (explicit NO_x) } \\ \text{ yield similar results). When } \Delta^{17}O(HO_2) \text{ is computed } \\ \text{ explicity, } \overline{\Delta^{17}O_{HNO_4 hydrol} (NO_3^-)} \text{ values increase moderately by ca. } 0.6\%. }$
- −  $\overline{\Delta^{17}O_{NO_3+RH}(NO_3^-)}$  and  $\overline{\Delta^{17}O_{N_2O_5hydrol}(NO_3^-)}$  are on the order of 40 and 33.5‰, respectively, and show a significant different between Case 1 (PSS NO<sub>x</sub> and Case 2 (explicit NO<sub>x</sub>), i.e. whether PSS is used to compute the  $\Delta^{17}O$  of NO<sub>x</sub> and NO<sub>y</sub>. Using the PSS formulation for NO<sub>x</sub> yields an overestimation on the order of 2‰ of the DIIS values for both these nitrate production channels.
- $\begin{array}{l} \overline{\Delta^{17}O_{HO_2+HO_2}\left(H_2O_2\right)} \mbox{ values are 0 for both Cases 1} \\ \mbox{and 2 (PSS and explicit NO_x, respectively), consistent with the fact that <math display="inline">\Delta^{17}O(HO_2)$  is assigned a value of 0. When  $\Delta^{17}O(HO_2)$  is computed explicitly, the  $\overline{\Delta^{17}O_{HO_2+HO_2}\left(H_2O_2\right)}$  reaches 1.1‰ (Case 3, explicit NO\_x and HO\_x). Taking into account non-mass dependent fractionation (NMDF) of 1‰ in the H+O\_2 reaction leads to increasing the  $\overline{\Delta^{17}O_{HO_2+HO_2}\left(H_2O_2\right)}$  value by 0.7‰.
- under the environmental conditions tested, whether isotopic exchange between OH and H<sub>2</sub>O is considered has no significant impact on all the reaction pathways considered. This simply indicates that  $\Delta^{17}O(OH)=0\%$  under these environmental conditions.

 Isotopic scrambling during the thermal and photolytical decomposition of and HNO<sub>4</sub> leads to lowering the DIIS value of all reaction pathways, except for H<sub>2</sub>O<sub>2</sub> production.

#### 4.1.3 Detailed analysis of the main results

#### Representation of diurnal variations of $\Delta^{17}O(NO_2)$

As shown on Fig. 4a,  $\Delta^{17}O(NO_2)$  exhibits diurnal variations with a maximum during the night and a minimum during the day, consistent with previous expectations (Morin, 2009). Figure 5 compares the results obtained using permanent photochemical steady-state (Case 1, NO<sub>x</sub> PSS) and explicitly computed (Case 2, NO<sub>x</sub> explicit, see Sect. 3.2.2). During the day, both calculations show a minimum at noon, on the order of 28‰. This is explained by the fact that during the day, the contribution of the NO+RO2 and NO+HO2 to the production of NO<sub>2</sub> peaks at noon, when peroxy radicals reach their maximum values (see Fig. 3). Owing to the short lifetime of NO<sub>2</sub> during the day, the result of the computation based on PSS is fully consistent with the explicit computation. This explains why  $\Delta^{17}O_{OH+NO_2}(NO_3^-)$  is the same under Case 1 and Case 2 (PSS and explicit NO<sub>x</sub>, respectively, see Table 5), because this pathway operates only during the day, when  $\Delta^{17}O(NO_2)$  has the same value in both cases.

The major difference between the two simulations occurs at night. Indeed, while the result from PSS leads  $\Delta^{17}O(NO_2)$ to reach values above 41‰ at night (i.e., on the order of  $\Delta^{17}O_{NO+O_3}(NO_2)$ ), the result from the explicit calculation does not exceed 39‰ at night, except for a limited period of time at dawn. The explanation for this behavior follows: at dusk, the  $\Delta^{17}O(NO_2)$  is fixed by the PSS conditions which prevail just before PSS recycling of NOx becomes insignificant. At this point, NO2 becomes relatively inert and its  $\Delta^{17}$ O does not vary anymore. As evidenced by Fig. 5, nighttime  $\Delta^{17}O(NO_2)$  corresponds to the  $\Delta^{17}O(NO_2)$  value computed at PSS when the lifetime of NO<sub>2</sub> is on the order of 10 min. It is then "frozen" until the dawn comes, along with the restart of photochemical activity. The difference between the two calculations lies between 2 and 3‰ under the conditions of the base model run. This explains why DIIS values for the



**Fig. 4.** Simulated diurnal variations of the isotopic signature  $\Delta^{17}O_i(nitrate)$  and  $\Delta^{17}O_i(H_2O_2)$  (**a**–**c**) and the strength (**d**–**f**) of the OH+NO<sub>2</sub>, NO<sub>3</sub>+RH nitrate production channel and H<sub>2</sub>O<sub>2</sub> production channel, respectively. Note that in the case of the OH+NO<sub>2</sub> reaction, the  $\Delta^{17}O_i$  values of individuals precursors (OH and NO<sub>2</sub>) are shown along with the isotopic signature of this reaction channel (see Sect. 2.4.1). Results originate from the base model run (T=293 K, 45° N, see Sect. 3.4 for details) under the isotopic hypotheses of Case 3 (explicit NO<sub>x</sub> and HO<sub>x</sub>). Hatched areas represent the period of the day when the considered reaction pathway proceeds insignificantly, thus the corresponding  $\Delta^{17}O$  which could then be transferred to the reaction products is irrelevant. The diurnally-integrated isotopic source (DIIS) values correspond to the average of values from the upper panel, weighted by the values of the lower one. Note that the strength of the production pathways are expressed in arbitrary units different in each cases, thus the panels (d) to (f) cannot be directly compared.

| Table 6.   | urnally-integrated source isotopic signature (DIIS) for the four nitrate and $H_2O_2$ production channels considered (in $\infty$ units) |
|------------|--|
| Overview   | the results from three model runs carried out at different periods of the year (W=1 January, Sp.=1 April and Su.=1 July                  |
| respective | , under mid-latitude (45° N) remote boundary layer conditions.   |

|   | (    | OH+NO <sub>2</sub> N |      | NO <sub>3</sub> +RH |      | N    | N2O5 hydrol |      | HNO <sub>4</sub> hydrol |      | H <sub>2</sub> O <sub>2</sub> |      |     |     |     |
|---|------|----------------------|------|---------------------|------|------|-------------|------|-------------------------|------|-------------------------------|------|-----|-----|-----|
|   | W    | Sp.                  | Su.  | W                   | Sp.  | Su.  | W           | Sp.  | Su.                     | W    | Sp.                           | Su.  | W   | Sp. | Su. |
| Case $1 - NO_x PSS$                     | 24.9 | 20.3                 | 17.3 | 42.7                | 42.6 | 42.3 | 35.2        | 35.1 | 34.9                    | 18.8 | 15.7                          | 13.6 | 0.0 | 0.0 | 0.0 |
| Case $2 - NO_x$ explicit                | 24.9 | 20.4                 | 17.3 | 40.7                | 40.2 | 39.8 | 33.8        | 33.2 | 32.8                    | 18.8 | 15.7                          | 13.6 | 0.0 | 0.0 | 0.0 |
| Case $3 - HO_x$ explicit                | 24.9 | 20.5                 | 17.4 | 40.8                | 40.2 | 39.8 | 33.8        | 33.3 | 32.8                    | 19.4 | 16.3                          | 14.2 | 1.1 | 1.1 | 1.0 |
| Case $4 - MIF$ in $H + O_2$             | 25.0 | 20.5                 | 17.5 | 40.8                | 40.2 | 39.8 | 33.8        | 33.3 | 32.8                    | 19.8 | 16.7                          | 14.6 | 1.8 | 1.8 | 1.7 |
| Case 5 – Scrambling of HNO <sub>4</sub> | 24.2 | 19.7                 | 16.7 | 38.0                | 38.0 | 38.0 | 31.4        | 31.2 | 31.0                    | 19.9 | 16.0                          | 13.7 | 2.9 | 1.5 | 1.1 |
| Case $6 - OH + H_2O$ isotopic exhange   | 25.1 | 20.5                 | 17.5 | 40.8                | 40.2 | 39.8 | 33.8        | 33.3 | 32.8                    | 19.4 | 16.3                          | 14.2 | 1.1 | 1.1 | 1.0 |

nighttime nitrate production pathways depend strongly on the method chosen to compute  $\Delta^{17}O(NO_x)$  and  $\Delta^{17}O(NO_y)$ , because much of the difference occurs at night.

#### $\Delta^{17}O(HO_2)$ and $\Delta^{17}O(H_2O_2)$

In Cases 3, 4, 5 and 6, where  $\Delta^{17}O(HO_2)$  is computed explicitly, non-zero  $\Delta^{17}O(HO_2)$  values are found. Figure 6 shows the diurnal variations of  $\Delta^{17}O(HO_2)$  and  $\Delta^{17}O(H_2O_2)$ . The non-zero  $\Delta^{17}O$  value in Case 3 (explicit NO<sub>x</sub> and HO<sub>x</sub>, on the order of 1‰, stems directly from the OH+O<sub>3</sub> reaction producing HO<sub>2</sub> with a non-zero  $\Delta^{17}O$  value, which is then mixed with other sources of HO<sub>2</sub>. The addition of non-mass dependent fractionation through the H+O<sub>2</sub> reaction, which is the dominant HO<sub>2</sub> production reaction, results in elevating the  $\Delta^{17}O(HO_2)$  value by roughly the magnitude of the isotopic fractionation constant. Combining the explicit calculation of the time evolution of  $\Delta^{17}O(HO_2)$  with the inclusion of non-mass dependent fractionation occurring during the H+O<sub>2</sub> reaction leads to daytime  $\Delta^{17}O(HO_2)$  values on the order of 2‰.

We note that the corresponding  $\Delta^{17}O_{HO_2+HO_2}$  (H<sub>2</sub>O<sub>2</sub>) values for either Case 3 or 4 (explicit NO<sub>x</sub> and HO<sub>x</sub> and NMDF H+O<sub>2</sub>, respectively), 1.1 and 1.8‰, respectively, are consistent with the experimental results of Savarino and Thiemens (1999a), who measured rainwater  $\Delta^{17}O(H_2O_2)$ 



**Fig. 5.** Diurnal variation of  $\Delta^{17}O(NO_2)$  calculated using two different approaches, i.e. permanent photochemical steady-state for NO<sub>x</sub> (Case 1, solid line) and the explicit computation of  $\Delta^{17}O(NO_x)$  (Case 2, dashed line). The dashed line representing the results of the simulations under Case 2 is overlaid with a color code exhibiting the diurnal variation of the photolytical lifetime of NO<sub>2</sub>. All values above 10 min are shown in red (at night, the photochemical lifetime of NO<sub>2</sub> is virtually infinite).

values ranging between 0.9 and 2.0‰, under coastal conditions in California.

### Impact of isotopic scrambling during the thermal and photolytical decomposition of HNO<sub>4</sub>

Case 5 tests the hypothesis where thermal decomposition and photolysis leads to isotopic scrambling between O atoms in molecules making up HNO<sub>4</sub> (see Table 4). Under such conditions, it is observed that the  $\Delta^{17}$ O of HO<sub>2</sub> increases, while  $\Delta^{17}$ O of NO<sub>x</sub> generally decreases. This behavior is reflected in the DIIS values of the relevant reactions (see Table 5). Detailed investigation of the reasons for this result reveals that much of the effect proceeds through the thermal decomposition of HNO<sub>4</sub>, especially at dusk when HNO<sub>4</sub> thermal decomposition is on the order of its formation rate due to reduced photochemical activity lowering the amount of HO<sub>2</sub>. Through slow but steady cycles of formation/decomposition, HNO<sub>4</sub> temporarily bridges the pool of oxygen atoms within  $NO_x$  and  $HO_x$ , leading to lowering the  $\Delta^{17}O$  of  $NO_x$  and increasing the  $\Delta^{17}$ O of HO<sub>x</sub> in a significant manner (see Fig. 6). In the following, we do not further discuss the impact of the hypothesis of Case 5, although model results are also given for this case.



**Fig. 6.** Simulated diurnal variations of  $\Delta^{17}O(H_2O_2)$  (top) and  $\Delta^{17}O(HO_2)$  (bottom) under the various hypotheses tested, from the results of the base model run. Note that in Case 1 and 2 (PSS and explicit NO<sub>x</sub>),  $\Delta^{17}O(HO_2)$  is set to 0. The hatched area covers night-time periods when the presence of HO<sub>2</sub> in the atmosphere is insignificant, hence the relevance of its  $\Delta^{17}O$  value.

#### 4.2 Sensitivity to atmospheric conditions

This section presents the results obtained under different conditions than in the base model run. For the sake of brevity, and since the physico-chemical reasons behind the observed behavior are similar to the phenomena described above, we focus our attention on DIIS values, which provide an efficient metric to compare model runs carried out under different environmental conditions.

#### 4.2.1 Impact of seasonal variations

Using the same chemical mechanism and the same initial chemical composition of the boundary layer, model runs were performed at different times of the year, i.e. starting from 1 January with a temperature of 283 K and 1 July with a temperature of 303 K to see whether seasonal variations in environmental conditions (temperature and actinic flux) can modify the conclusions reached above for the springtime model run, started on 1 April with a temperature of 293 K. The photochemical activity increases monotonically from the winter to summer model runs, as expected (chemical data not shown).

The model shows that the DIIS of daytime nitrate production channels  $(\overline{\Delta^{17}O_{OH+NO_2}(NO_3^-)})$  and  $\overline{\Delta^{17}O_{HNO_4}hydrol}(NO_3^-))$  are most dependent on the season, due to their strong ties to photochemical activity, which controls  $\Delta^{17}O$  of NO<sub>2</sub> and HNO<sub>4</sub> through photochemical steady state during the day. The  $\overline{\Delta^{17}O_{OH+NO_2}(NO_3^-)}$  varies

|   | $OH+NO_2$ |       | NO <sub>3</sub> +RH |       | N <sub>2</sub> O <sub>5</sub> hydrol |       | HNO <sub>4</sub> hydrol |       | $H_2O_2$ |       |
|---|-----------|-------|---------------------|-------|--------------------------------------|-------|-------------------------|-------|----------|-------|
|   | 45° N     | 80° N | 45° N               | 80° N | 45° N                                | 80° N | 45° N                   | 80° N | 45° N    | 80° N |
| Case 1 – NO <sub>x</sub> PSS            | 20.3      | 26.0  | 42.6                | 42.1  | 35.1                                 | 34.4  | 15.7                    | 19.6  | 0.0      | 0.0   |
| Case $2 - NO_x$ explicit                | 20.4      | 25.9  | 40.2                | 41.1  | 33.2                                 | 33.6  | 15.7                    | 19.5  | 0.0      | 0.0   |
| Case $3 - HO_x$ explicit                | 20.5      | 25.9  | 40.2                | 41.1  | 33.3                                 | 33.6  | 16.3                    | 19.9  | 1.1      | 0.8   |
| Case 4 – MIF in $H + O_2$               | 20.5      | 26.0  | 40.2                | 41.1  | 33.3                                 | 33.7  | 16.7                    | 20.3  | 1.8      | 1.6   |
| Case 5 – Scrambling of HNO <sub>4</sub> | 19.7      | 24.9  | 38.0                | 36.2  | 31.2                                 | 28.5  | 16.0                    | 21.1  | 1.5      | 3.8   |
| Case $6 - OH + H_2O$ isotopic exhange   | 20.5      | 28.1  | 40.2                | 41.1  | 33.3                                 | 33.6  | 16.3                    | 20.0  | 1.1      | 0.9   |

**Table 7.** Comparison of diurnally-integrated source isotopic signature (DIIS, in % units) for the four nitrate and H<sub>2</sub>O<sub>2</sub> production channels considered under mid-latitude (45° N, 293 K) and Arctic (80° N, 253 K) conditions in springtime (1 April).

significantly, from ca. 17‰ in summer to ca. 25‰ during the winter.  $\overline{\Delta^{17}O_{HNO_4hydrol}(NO_3^-)}$  also varies, from ca. 14‰ in summer to ca. 19‰ in winter.

In contrast, the DIIS of nighttime nitrate production channels exhibits a stronger dependence upon the isotopic assumption (as detailed in Sect. 4.1.3), but shows little seasonal variations. The biggest variation is between Case 1 (PSS NO<sub>x</sub>) and Case 2 (explicit NO<sub>x</sub>), the former leading to an overestimation ranging from 1.5 to 2.0‰, from winter to summer for both N<sub>2</sub>O<sub>5</sub> hydrolysis and NO<sub>3</sub>+RH. The  $\overline{\Delta^{17}O_{NO_3+RH}(NO_3^-)}$  and  $\overline{\Delta^{17}O_{N_2O_5hydrol}(NO_3^-)}$  both vary within about 1‰ seasonally, and remain on the order of 40 and 33‰ year-round, respectively.

Last,  $\overline{\Delta^{17}O_{HO_2+HO_2}(H_2O_2)}$  values show little seasonal variation (less than 0.1‰). They remain consistently on the order of 1‰ when only the OH+O<sub>3</sub> reaction is responsible for  $\Delta^{17}O$  transfer from O<sub>3</sub> to HO<sub>2</sub>. They increase to around 1.7‰ when non-mass dependent fractionation of 1‰ is considered throughout the reaction H+O<sub>2</sub>  $\rightarrow$  HO<sub>2</sub>. The only major change occurs in Case 5, i.e. considering scrambling during the thermal decomposition and photolysis of HNO<sub>4</sub>.

Under the conditions experienced for our base model run, the isotopic exchange reaction between OH and H<sub>2</sub>O does not lead to  $\Delta^{17}$ O(OH) values significantly different from 0. This explains why the DIIS values for Case 6 (OH+H<sub>2</sub>O isotopic exchange) are very similar to that of Case 3 (explicit NO<sub>x</sub> and HO<sub>x</sub>). This effects becomes significant only at lower temperatures (see Morin et al., 2007) and is further explored in Sect. 4.2.2.

#### 4.2.2 Higher latitude and colder conditions

A simulation was carried out under springtime Arctic conditions, i.e. a latitude of 80° N and temperature of 253 K, starting from 1 April. The results of this comparison is shown in terms of DIIS values in Table 7.

The DIIS values of daytime nitrate production channels show a strong difference between Arctic and mid-latitude conditions. With the exception of Case 6 (OH+H<sub>2</sub>O isotopic exchange), the Arctic  $\overline{\Delta^{17}O_{OH+NO_2}(NO_3^-)}$  values are ca. 5.5‰ higher than at mid-latitudes, which simply stems from reduced photochemical recycling under reduced insulation prevailing in the Arctic and colder temperatures. It is noteworthy that under Arctic conditions, owing to the lower temperatures prevailing, the value of  $\overline{\Delta^{17}O_{OH+NO_2}(NO_3^-)}$  is 2‰ higher under Case 6 (OH+H<sub>2</sub>O isotopic exchange) than under other cases (except Case 5). The diurnal variations of  $\Delta^{17}O(NO_2)$  is similar for Case 2 (explicit NO<sub>x</sub>) and Case 6 (OH+H<sub>2</sub>O isotopic exchange) under Arctic conditions, demonstrating that all of the difference observed stems from the fact that  $\Delta^{17}O(OH)$  amounts ca. 6‰ under Arctic conditions and Case 6, consistent with the initial estimates provided by Morin et al. (2007).

 $\label{eq:constraint} \begin{array}{|c|c|c|c|c|} \hline The values of $\overline{\Delta^{17}O_{NO_3+RH}\left(NO_3^{-}\right)}$ and $\overline{\Delta^{17}O_{N_2O_5\,hydrol}\left(NO_3^{-}\right)}$ are very similar under mid-latitude and Arctic conditions, on the order of 40.5‰ and 33.5‰, respectively. Note also that the impact of the hypothesis of Case 6 (OH+H_2O isotopic exchange) on the DIIS values is insignificant. Arctic values of $\overline{\Delta^{17}O_{HO_2+HO_2}\left(H_2O_2\right)$ tend to be a little lower by a few tenths of ‰. The effect of Case 6 is limited to 0.1‰ in terms of $\overline{\Delta^{17}O_{HO_2+HO_2}\left(H_2O_2\right)$.} \end{array}$ 

In summary, the impact of colder and more boreal environmental conditions is mostly seen for the DIIS of daytime nitrate production, upon which photochemical conditions and the  $\Delta^{17}$ O of OH have a direct impact. The DIIS of nighttime nitrate production channels as well as H<sub>2</sub>O<sub>2</sub> seem to be fairly insensitive to these factors.

#### 4.2.3 Higher initial NO<sub>x</sub> mixing ratio

A further simulation was carried out under springtime midlatitude conditions (45° N, 293 K), with an initial NO<sub>x</sub> mixing ratio of 2 nmol mol<sup>-1</sup> instead of 20 pmol mol<sup>-1</sup> in the base run. The results in terms of DIIS are shown in Table 8. The impact of performing the PSS approximation to compute  $\Delta^{17}$ O of NO<sub>x</sub> is similar, i.e. there is no impact on daytime nitrate production pathways and H<sub>2</sub>O<sub>2</sub>. It is more visible in the case of nighttime nitrate production pathways, although the difference between Case 1 (PSS NO<sub>x</sub>) and Case 2 (explicit NO<sub>x</sub>) for nighttime nitrate production pathways is smaller than under the conditions of the base model run.

**Table 8.** Comparison of diurnally-integrated source isotopic signature (DIIS, in % units) for the four nitrate and H<sub>2</sub>O<sub>2</sub> production channels considered, respectively under mid-latitude (45° N, 293 K) and under starting NO<sub>x</sub> levels of 20 pmol mol<sup>-1</sup> (base) and 2 nmol mol<sup>-1</sup> (high NO<sub>x</sub>) in springtime (1 April).

|   | OH+NO <sub>2</sub> |                         | NO <sub>3</sub> +RH |                         | N <sub>2</sub> O <sub>5</sub> hydrol |                         | HNO <sub>4</sub> hydrol |                         | H <sub>2</sub> O <sub>2</sub> |                         |
|---|--------------------|-------------------------|---------------------|-------------------------|--------------------------------------|-------------------------|-------------------------|-------------------------|-------------------------------|-------------------------|
|   | base               | high<br>NO <sub>x</sub> | base                | high<br>NO <sub>x</sub> | base                                 | high<br>NO <sub>x</sub> | base                    | high<br>NO <sub>X</sub> | base                          | high<br>NO <sub>x</sub> |
| Case 1 – NO <sub>x</sub> PSS            | 20.3               | 17.3                    | 42.6                | 42.1                    | 35.1                                 | 34.4                    | 15.7                    | 17.5                    | 0.0                           | 0.0                     |
| Case $2 - NO_x$ explicit                | 20.4               | 17.3                    | 40.2                | 41.1                    | 33.2                                 | 33.6                    | 15.7                    | 17.5                    | 0.0                           | 0.0                     |
| Case $3 - HO_x$ explicit                | 20.5               | 17.4                    | 40.2                | 41.1                    | 33.3                                 | 33.6                    | 16.3                    | 18.2                    | 1.1                           | 1.3                     |
| Case 4 – MIF in $H + O_2$               | 20.5               | 17.5                    | 40.2                | 41.1                    | 33.3                                 | 33.7                    | 16.7                    | 18.5                    | 1.8                           | 1.9                     |
| Case 5 – Scrambling of HNO <sub>4</sub> | 19.7               | 16.7                    | 38.0                | 36.2                    | 31.2                                 | 28.5                    | 16.0                    | 19.7                    | 1.5                           | 4.5                     |
| Case $6 - OH + H_2O$ isotopic exhange   | 20.5               | 17.5                    | 40.2                | 41.1                    | 33.3                                 | 33.6                    | 16.3                    | 18.2                    | 1.1                           | 1.3                     |

**Table 9.** Comparison between the DIIS values for  $OH+NO_2$  and  $NO_3+RH$  with the corresponding values assigned to these reactions by Alexander et al. (2009), in % units. The values deduced from the algorithm presented by Alexander et al. (2009) are compared to DIIS values computed using Cases 1 and 3, under the conditions of the base model run (mid-latitudes).

|                                     | OH         | $+NO_2$                    | NO         | 3+RH                       |
|-------------------------------------|------------|----------------------------|------------|----------------------------|
|                                     | this study | Alexander<br>et al. (2009) | this study | Alexander<br>et al. (2009) |
| Case 1 – NO <sub>x</sub> PSS        | 20.3       | 19.3                       | 42.6       | 42.6                       |
| Case $3 - HO_x$ and $NO_x$ explicit | 20.5       |                            | 40.2       |                            |

Due to enhanced photochemical activity fueled by higher initial NO<sub>x</sub> levels, the  $\overline{\Delta^{17}O_{OH+NO_2}(NO_3^-)}$  values are reduced by 3‰ when higher initial NO<sub>x</sub> levels are set. The impact on nighttime nitrate production channels is limited, with  $\overline{\Delta^{17}O_{NO_3+RH}(NO_3^-)}$  and  $\overline{\Delta^{17}O_{N_2O_5hydrol}(NO_3^-)}$  values on the order of 40.5‰ and 34‰, respectively, varying about 1‰ across this strong chemical gradient.

#### 5 Discussion and implications

#### 5.1 Implications for modeling $\Delta^{17}O(NO_3^-)$

We compare here our results to the implementation of  $\Delta^{17}O(NO_3^-)$  into the GEOS-Chem chemistry transport model, which was recently carried out by Alexander et al. (2009). In this work,  $\Delta^{17}O(NO_x)$  was computed under the hypothesis of photochemical steady-state. For the sole daytime nitrate production channel considered (OH+NO<sub>2</sub>),  $\Delta^{17}O(NO_2)$  was estimated using the  $\alpha$  value (see Sect. 3.2.1) computed using accumulated reaction rates between 10:00 and 14:00 solar time. For nighttime nitrate production channels, Alexander et al. (2009) used the photochemical steadystate formalism using NO<sub>2</sub> production rates accumulated between 00:00 and 02:00 solar time. We compute the  $\Delta^{17}O$ inherited by atmospheric nitrate through the OH+NO<sub>2</sub> and NO<sub>3</sub>+RH using the algorithm of Alexander et al. (2009) presented above, and compare it to the corresponding DIIS values. The results are given in Table 9. We find that the algorithm introduced by Alexander et al. (2009) underestimates the isotopic signature of the OH+NO<sub>2</sub> channel by 1‰ under springtime conditions, because it ignores contributions of this channel before 10:00 and after 14:00, when  $\Delta^{17}O(NO_2)$  is relatively higher than during noontime but the OH+NO<sub>2</sub> reaction proceeds significantly. The underestimates ranges between 0.5 and 1.5‰ when the comparison extends to other seasons and latitudes considered in the model runs. Alexander et al. (2009) also overestimate the isotopic signature of the NO<sub>3</sub>+RH channel by 1.6‰, due to the fact that they use PSS equations to derive  $\Delta^{17}O(NO_y)$  at night, which has been proven above to cause significant overestimation of the DIIS of nighttime nitrate production channels.

We strongly suggest that DIIS values are implemented in large-scale modeling frameworks, such as GEOS-Chem, to avoid performing such errors. To facilitate this, we provide below a method to implement this concept within a model such as GEOS-Chem, using the formalism of Alexander et al. (2009). The method is independent of the choice of  $\Delta^{17}O(O_3)$  although they are based on the hypothesis that  $\Delta^{17}O(O_3)$  remains constant throughout the day. Indeed, as an alternative to DIIS values introduced in Sect. 2.4.1 above, we introduce here the concept of *effective*  $\alpha$  values, as follows. We combine Eqs. (6), (7) and (8) with the  $\Delta^{17}O$  transfer functions under photochemical steady-state conditions, described in Sect. 3.2.1. During the day, at PSS, the isotopic signature of NO<sub>2</sub>+OH, NO<sub>3</sub>+RH, N<sub>2</sub>O<sub>5</sub> hydrolysis, respectively, are given by:

$$\Delta^{17}O_{\rm NO_2+OH}(\rm NO_3^-) = 2/3\alpha \,\Delta^{17}O_{\rm NO+O_3}(\rm NO_2)$$
(17)

$$\Delta^{17}O_{NO_3+RH}(NO_3^-) = 1/3 \left( 2\alpha \Delta^{17}O_{NO+O_3}(NO_2) + \Delta^{17}O(O_3^*) \right)$$
(18)

$$\Delta^{17}O_{N_2O_5 hydrol} \left( NO_3^{-} \right) = 1/6 \left( 4\alpha \Delta^{17}O_{NO+O_3} (NO_2) + \Delta^{17}O(O_3^{\star}) \right)$$
(19)

These three equations can be generalized as follow: rather than computing the instantaneous isotopic signature of each nitrate source considered, under PSS conditions, let us examine their diurnally integrated isotopic signature. It is possible to compute the  $\alpha$  value matching the DIIS values explicitly computed, using the formalism of Eqs. (17), (18) and (19):

$$\overline{\Delta^{17}O_{\text{OH}+\text{NO}_2}(\text{NO}_3^-)} = 2/3\overline{\alpha_{\text{NO}_2+\text{OH}}}\Delta^{17}O_{\text{NO}+\text{O}_3}(\text{NO}_2)$$
(20)

$$\overline{\Delta^{17}O_{\text{NO}_3+\text{RH}}(\text{NO}_3^-)} = 1/3 \left( 2\overline{\alpha_{\text{NO}_3+\text{RH}}} \Delta^{17}O_{\text{NO}+\text{O}_3}(\text{NO}_2) + \Delta^{17}O(\text{O}_3^*) \right) \quad (21)$$

$$\overline{\Delta^{17}O_{N_2O_5hydrol}(NO_3^-)} = 1/6 \left( 4\overline{\alpha_{N_2O_5hydrol}} \Delta^{17}O_{NO+O_3}(NO_2) + \Delta^{17}O(O_3^{\star}) \right)$$
(22)

The then defined *effective*  $\alpha$ , denoted  $\overline{\alpha}$ , are specific to each nitrate production pathways, and can serve to expand the formalism defined at PSS to account for the fact that PSS does not hold to quantitatively compute the isotopic signature of atmospheric nitrate sources. Such effective  $\alpha$  values can then be implemented in a model like GEOS-Chem (Alexander et al., 2009). Indeed, in this model implementation, the isotopic signatures of nitrate production channels were computed at a daily time resolution (B. Alexander, personal communication 2011).  $\overline{\alpha_{NO_2+OH}}$  is conceptually equivalent to  $\alpha_{dav}$  as defined in Alexander et al. (2009) because this pathway proceeds during the day. Likewise, Alexander et al. (2009) defined a  $\alpha_{night}$  value which holds for N<sub>2</sub>O<sub>5</sub> hydrolysis and the  $NO_3 + RH$  reactions. In the cases tested, our study indicated that  $\overline{\alpha_{NO_3+RH}}$  has the same value as  $\overline{\alpha_{N_2O_5hvdrol}}$ , justifying the use of a single  $\overline{\alpha}$  value for "nighttime" nitrate production channels. The  $\overline{\alpha}$  values extracted from data in Table 9 for the  $OH + NO_2$  channel are equal to 0.73 and 0.69 for the explicit calculation and Alexander et al. (2009), respectively. For the NO<sub>3</sub>+RH channel, the  $\overline{\alpha}$  values are equal to 0.93 and 0.99, respectively. The errors in terms of DIIS, as detailed above, translate in errors in terms of  $\overline{\alpha}$ .

In practice, the issue associated with the computation of  $\overline{\alpha_{NO_2+OH}}$  in GEOS-Chem can easily be solved as follows: rather than computing  $\alpha$  values over an arbitrary time window which causes a systematic underestimation at least on the order of 0.04 in terms of  $\overline{\alpha}$ , we recommend to extend the temporal range of integration to the full length of the day (i.e., from 00:00 to 24:00 solar time), and scale the instantaneous PSS  $\alpha$  values with the rate of reaction OH+NO<sub>2</sub>. Indeed, we have shown that  $\overline{\Delta^{17}O_{OH+NO_2}(NO_3^-)}$  values computed using PSS approximations match the DIIS values obtained through the explicit modeling of  $\Delta^{17}O$  of NO<sub>x</sub> and NO<sub>y</sub>. The issue with the nighttime nitrate production channels should be solved without resorting to using photochemical steady state equations at night, since we have shown that this leads to systematically erroneous results and is based on a scientific oxymoron. The model results show that  $\Delta^{17}O_{NO_3+RH}(NO_3^-)$  and  $\Delta^{17}O_{N_2O_5\,hydrol}(NO_3^-)$  do not vary significantly over night. This behavior stems from the fact that the main driver of change of  $\Delta^{17}$ O of nitrate precursors, that is photochemical cycling between NO<sub>x</sub>, HO<sub>x</sub> and ozone is inactive at night, thereby "freezing" the  $\Delta^{17}$ O signature of most reactants until dawn. Given the low sensitivity of DIIS values, and thus  $\overline{\alpha}$  values, for nighttime nitrate production channels to environmental conditions such as temperature, actinic flux and NO<sub>x</sub> levels, a conservative approach may be to use a fixed value of 0.94 for the  $\overline{\alpha}$  of NO<sub>3</sub>+RH and N<sub>2</sub>O<sub>5</sub> hydrolysis, respectively. Indeed, this value corresponds to the middle of the range covered (within 0.02, minimum in summer and maximum in winter) by our model runs, spanning a wide range of atmospheric and environmental conditions.

#### 5.2 Implication for $\Delta^{17}O(H_2O_2)$

An interesting implication of our work is the fact that under all tested environmental conditions, the model predicts non-zero  $\Delta^{17}O(H_2O_2)$  even without invoking non-mass dependent fractionation through the H+O<sub>2</sub> reaction, as evidenced by Savarino and Thiemens (1999a). This significant  $\Delta^{17}O(H_2O_2)$ , on the order of 1‰, stems from the OH+O<sub>3</sub> reaction and could be used in the future to probe the level of photochemical activity of a given air parcel through measurements of  $\Delta^{17}O(H_2O_2)$  either in the gas-phase or in rainwater, as suggested by Savarino and Thiemens (1999b).

#### 5.3 Open questions

Two untested assumptions of this work are the hypothesis that  $\Delta^{17}O(O_3)$  remains constant throughout the day, and that the thermal decomposition and photolysis of HNO<sub>4</sub> does not lead to any isotopic scrambling between the two molecules making up these dimers. The assumption related to the diurnal variations of  $\Delta^{17}O(O_3)$  can now be addressed using field measurements carried out using a chemical probing method based upon the  $NO_2^- + O_3$  reaction (Michalski and Bhattacharya, 2009; Vicars et al., 2011). As indicated above, should  $\Delta^{17}O(O_3)$  exhibit significant diurnal variations, several quantitative results provided above for illustration purposes would become obsolete - note that the same would apply to all model studies carried out so far, which all assume diurnally invariant  $\Delta^{17}O(O_3)$  (e.g. Michalski et al., 2003; Morin et al., 2008; Kunasek et al., 2008; Dominguez et al., 2009; Michalski and Xu, 2010; Alexander et al., 2009). Nevertheless, the methodology presented here would remain valid and the developed framework is already suited to account for diurnal variations in  $\Delta^{17}O(O_3)$ (Vicars et al., 2011). The question related to the chemical mechanism operating during thermal decomposition requires advanced chemical physics modeling at the molecular scale.

Alternatively, studies on  $\Delta^{17}O(H_2O_2)$  may partly solve this issue, since we have shown that  $\overline{\Delta^{17}O_{HO_2+HO_2}(H_2O_2)}$  is most sensitive to this assumption.

Additional sources of uncertainties have not been addressed in detail in this study, such as the impact of nonzero  $\Delta^{17}O(O_2)$  and  $\Delta^{17}O(H_2O)$  on  $\Delta^{17}O$  of reactive and secondary species: although measurable (Barkan and Luz, 2003, 2005), the  $\Delta^{17}O$  of  $O_2$  and  $H_2O$  are expected to be small and should affect negligibly the discussion presented above. We have also chosen to not propagate experimental and theoretical uncertainties associated with the  $\Delta^{17}O$  transfer rate of  $\Delta^{17}O(O_3)$  along with chemical reactions mostly for clarity reasons, to rather focus on the overall methodological aspects of the explicit modeling of the  $\Delta^{17}O$  transfer through a full photochemical mechanism.

#### 6 Conclusions

This study addresses in detail the question of the impact of diurnal variations of  $\Delta^{17}O$  of short-lived reactive species on secondary species such as atmospheric nitrate and H<sub>2</sub>O<sub>2</sub>. Using a state of the art photochemical box model, the time evolution of  $\Delta^{17}O$  of NO<sub>x</sub>, NO<sub>y</sub> and HO<sub>x</sub> is computed under various sets of hypotheses pertaining to the method of computing the  $\Delta^{17}O$  values, reflecting different levels of simplifying approximations. Most of the conclusions of this article are drawn from model simulations carried out under clean atmospheric conditions in mid-latitudes ; their broader relevance and robustness is however assessed using model simulations carried out under clean under a 100-fold increase in initial NO<sub>x</sub> mixing ratio.

The primary goal of this study was to demonstrate that using a detailed box-modeling study to assess the isotopic signature of various nitrate and H<sub>2</sub>O<sub>2</sub> production pathways is feasible and provides relevant information to larger-scale modeling studies. In the meantime, essential features of the coupling between chemical reactions and the  $\Delta^{17}$ O of key atmospheric species were described and are most likely also valid under different environmental contexts. Taking this study as an initial step, the model could rather easily be extended to account for gas/particles interactions, which are of primary importance for the budget of NO<sub>x</sub>, and to simulate  $\Delta^{17}$ O values of secondary species such as atmospheric nitrate and H<sub>2</sub>O<sub>2</sub> under atmospheric contexts as different as over continents (urban polluted, tropical etc.) and under polar conditions including more complex and realistic chemical mechanisms. We believe that the present study provides the necessary framework for carrying out this work under conditions that will make it usable by larger-scale modeling studies, or for the interpretation of short-term intensive measurement campaigns using a modeling tool analogous to CAABA/MECCA.

## Supplement related to this article is available online at: http://www.atmos-chem-phys.net/11/3653/2011/ acp-11-3653-2011-supplement.pdf.

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