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Stimulated infrared emission from rocks: assessing a stress indicator

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F. T. Freund, A. Takeuchi, B. W. S. Lau, A. Al-Manaseer, C. C. Fu, et al.. Stimulated infrared emission from rocks: assessing a stress indicator. *eEarth Discussions*, 2006, 1 (2), pp.97-121. hal-00298243

HAL Id: hal-00298243

<https://hal.science/hal-00298243>

Submitted on 18 Jun 2008

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Stimulated infrared emission from rocks

F. Freund et al.

Stimulated infrared emission from rocks: assessing a stress indicator

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Received: 16 May 2006 – Accepted: 3 July 2006 – Published: 24 July 2006

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Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

EGU

Abstract

To study the effect of stress-activated positive hole (p-hole) charge carriers on the infrared (IR) emission from rocks, we subjected a portion (~ 10 vol.%) of a large ($60 \times 30 \times 7.5 \text{ cm}^3$) block of anorthosite, a nearly monomineralic (Ca-rich feldspar) igneous rock, to uniaxial deviatoric stress up to failure. We measured the IR emission from a flat surface ≈ 40 cm from the stressed rock volume over the $800\text{--}1300 \text{ cm}^{-1}$ ($7.7\text{--}12.5 \mu\text{m}$) range. Upon loading, the intensity and spectrum of the IR emission change. Narrow bands near instantly appear at 930 cm^{-1} ($10.75 \mu\text{m}$), 880 cm^{-1} ($11.36 \mu\text{m}$), 820 cm^{-1} ($12.4 \mu\text{m}$) plus additional bands in the $1000\text{--}1300 \text{ cm}^{-1}$ ($10.0\text{--}7.7 \mu\text{m}$) range. Upon further loading the bands broaden and shift. Their intensities increase but also fluctuate. Near the emission maxima at 300 K, at 1150 cm^{-1} and 1030 cm^{-1} (8.7 and $9 \mu\text{m}$), barely any intensity increase occurs suggesting that the temperature of the surface does not actually increase. We propose that the observed narrow IR emission bands arise from vibrationally excited O-O stretching modes which form when p-hole charge carriers (activated in the stressed rock) spread into the unstressed portion of the rock to the surface, where they recombine and radiatively decay. The effect, stimulated IR emission due to hole-hole recombination, may help explain the enhanced IR emission seen in night-time satellite images of the land surface before major earthquakes known as “thermal anomalies”.

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

1 Introduction

Since the late 1980s and early 1990s, non-stationary areas of enhanced IR emission from the land surface have been recognized in night-time satellite images, which seem to be linked to impending earthquake activity and the build-up of stresses deep below (Gorny et al., 1988; Qiang et al., 1991; Qiang et al., 1990; Srivastav et al., 1997). The areas of enhanced IR emission are referred to as “thermal anomalies”. The reported excursions in the land surface temperature reach 2–4°C, occasionally higher.

These thermal anomalies have remained enigmatic (Cui et al., 1999; Srivastav et al., 1997; Tronin, 2000; Tronin, 2002; Tronin et al., 2004). In most cases the increase in radiative temperature seems not to correlate to meteorological ground data. The rapidity with which the temperature variations appear and disappear rules out that their cause is the flow of Joule heat from a source below. Several processes have been invoked to account for the apparent rise in temperature: (i) rising fluids that could lead to the emanation of warm gases (Gorny et al., 1998); (ii) rising well water levels and changing moisture contents in the soil (Chadha et al., 2003); (iii) diffuse CO₂ emanation causing a “local greenhouse” effect (Quing et al., 1991; Tronin, 1999; Tronin, 2002); (iv) Near-ground air ionization due to enhanced radon emission leading to the condensation of water vapor from the atmosphere and, hence, to the release of latent heat (Pulinets et al., 2006). However, upon closer inspection, none of these explanations seems to be able to adequately account for the characteristic features of the “thermal anomalies”, for the rapidity with which the large areas of enhanced IR emission appear and disappear, and for the often reported lack of correlation with ground-station meteorological data.

In this paper, we report on laboratory experiments conducted to test a fundamentally different hypothesis: that the enhanced IR emission from a rock surface may be controlled by electronic charge carriers that are activated within the rock through the application of deviatoric stress.

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

2 Experimental part

We chose anorthosite, an essentially monomineralic feldspar rock composed of the Ca-rich plagioclase labradorite. Our sample, available under the trade name “Blue Pearl”, from Larvik, Norway, is very coarse-grained with crystals up to 2-4 cm in size.

The density of the rock is 2.7 g/cm^3 and its compressive strength 181–187 MPa.

We uniaxially stressed a relatively small portion, 10–12 vol.%, of an air-dry anorthosite slab, $60 \times 30 \times 7.5 \text{ cm}^3$, via a pair of pistons (11.25 cm diameter), electrically insulated from the rock through 0.8 mm thick sheets of high density polyethylene with a resistivity of $>10^{14} \Omega \text{ cm}$. We applied the load off-center as sketched in Fig. 1, about four piston diameters ($>40 \text{ cm}$) away from the IR-emitting surface. The off-center loading concentrated the stresses in the back portion of the slab. Figure 1 displays the principal stress by finite element analysis. Under these conditions, the front face of the slab from where we record the IR emission remains essentially stress-free, especially during the early phase of loading when the absolute stress applied to the back part of the slab is still relatively small. We applied the load at a constant rate of 6.3 MPa/min up to failure, using a hydraulic SATEC press, model RD 2000 kN.

Emission spectra were recorded from a flat, circular area, 5 cm diameter, off the front face of the rock, smooth “as received”, i.e. cut with a diamond saw. We used a Bomen MB-100 FT-IR spectroradiometer equipped with a Peltier-cooled HgCdTe detector and two integrated blackbody emitters for internal calibration, one at ambient temperature, the other at 60°C , collecting the IR radiation sequentially from the sample and the two blackbody emitters. This spectroradiometer has been used extensively in laboratory and field calibration for NASA’s MODIS spectrometer currently flying on the TERRA and AQUA satellites (Li et al., 1999; Wan et al., 1996). During our measurement, the laboratory was semi-darkened. The space between the rock and the spectroradiometer, about 1 m, was shielded from ambient light. During the run, the movement of people in the room was restricted to avoid changes in the reflected IR radiation field.

The spectra were recorded over the wavenumber range $700\text{--}1400 \text{ cm}^{-1}$ (7.14--

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

14.25 μm) at 2 cm^{-1} steps. Each FT-IR file consists of 25 scans from the rock surface plus 5 scans from the ambient temperature and 5 scans from the 60°C blackbody emitters for temperature calibration. Each file took 40 s to acquire and store. The radiometric noise at the single scan level was about 100 mK, improving to about 50 mK upon averaging 25 scans for each file and to about 10 mK upon averaging 10 files. The run lasted 2200 s (36 min 40 s), of which the first 400 s (6 min 40 s) were recorded without applying a load. The IR flux emitted from the rock surface was calculated in temperature equivalents (energy equivalents kT) with reference to the built-in calibration emitter surfaces, using the same algorithm as for MODIS calibration runs.

3 Results

Figure 3 shows the pre-loading room temperature (300 K) IR emission spectrum of the anorthosite from 800 to 1300 cm^{-1} ($7.7\text{--}12.5\ \mu\text{m}$) averaged from the 250 pre-loading scans. The 300 K spectrum has two maxima, around 1020 and 1190 cm^{-1} (9.8 and $8.4\ \mu\text{m}$ respectively), plus a smaller emission peak around 1110 cm^{-1} ($9.0\ \mu\text{m}$). The emission bands are characteristic of Si–O and Al–O stretching modes thermally activated at 300 K (Johnson, et al., 2002). The room temperature emission is primarily due to downward transitions from the thermally populated first vibrationally excited states, i.e. from quantum number $n=1$ to their ground states $n=0$.

Figure 4 shows a 3-D plot of the intensity variations over the $7.4\text{--}14.3\ \mu\text{m}$ range (700 to 1350 cm^{-1}) as recorded during this run as a function of time before loading and during loading up to failure. The time axis is labeled in seconds. During the first 6 min 40 s, as we acquired 250 pre-load scans (10 files), the emitted IR intensity is constant. This is consistent with a stable ambient temperature environment in the laboratory. However, as soon as we applied the load, we see an overall increase in the emitted IR intensity. The increase is near-instant upon application of the load and new emission bands appear. Upon further increases of the load, we see that the intensities fluctuate strongly. Such fluctuations had already been observed during

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

earlier IR emission experiments (Freund et al., 2002b; Freund et al., 2003). The IR intensity fluctuations resemble current fluctuations seen during measurements of the “battery currents”, which flow out of a stressed rock volume through unstressed rock (Freund et al. 2006, 2004b).

The difference spectra in Fig. 5 show the excess intensity emitted over the spectral range as a function of load and time. The difference spectra were obtained by subtracting the average of the 10 pre-load files from each of the files acquired subsequently during loading up to failure. There are three outstanding features:

1. Narrow emission bands appear immediately upon loading, within 40 s, specifically three bands between $800\text{--}950\text{ cm}^{-1}$ ($10.5\text{--}12.5\text{ }\mu\text{m}$) plus a few narrow bands at higher wavenumbers (shorter wavelengths).
2. The intensity fluctuations as a function of time are synchronous across the entire spectral range, though the fluctuations of different bands are not evenly distributed.
3. The excess intensity emitted during loading is concentrated at wavelengths that are different from those of the two emission maxima in the pre-load emission spectrum.

This last point is highlighted in Fig. 6 where we superpose onto the pre-load 300 K spectrum the excess IR intensity integrated over the entire run. The narrow emission bands are preserved even after integration, in particular the three diagnostically significant bands in the $800\text{--}950\text{ cm}^{-1}$ ($10.5\text{--}12.5\text{ }\mu\text{m}$) range and a prominent narrow band around 1150 cm^{-1} ($8.7\text{ }\mu\text{m}$). While the pre-load spectrum exhibits intensity maxima at 1080 cm^{-1} and 1210 cm^{-1} (9.7 and $8.5\text{ }\mu\text{m}$), the excess emission has minima near these same wavenumbers/wavelengths.

During two earlier measurement sessions, we had the opportunity of record the IR emission from the surface of a other blocks of rock stressed a distance away from the emitting surface. We observed similar spectral features, in particular narrow emission

Stimulated infrared emission from rocksF. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

bands at the beginning of loading, which broadened with increasing stress (Freund et al., 2002a, 2003). For these earlier runs we had used both anorthosite and granite, a multi-mineralic rock with about 1/3 quartz, 1/3 plagioclase feldspar and 1/3 potassium feldspar. As reported excess IR was also emitted in these cases in narrow bands in the 800–950 cm⁻¹ (10.5–12.5 μm) range plus some broader bands in the 1150 cm⁻¹ (8.7 μm) range. The experiment presented here confirms these earlier observations, though the band positions were slightly different, probably due to the more complex mineralogy of the granite.

4 Discussion

Rock fracture experiments as described in the literature are typically performed under conditions emulating ASTM C170-50 and DIN 52102 procedures, i.e. with cylindrical test samples loaded over their entire cross section (Brady and Rowell, 1986; Lockner, 1993; Rowell et al., 1981; Warwick et al., 1982; Yoshida and Ogawa, 2004). Loading over the entire cross section, however, causes stress gradients in the surface and subsurface region. It causes the surface to bulge leading to tensile stresses, which cause microfracturing. Microfracturing leads to intense visible and IR emission over a broad spectral range (Brady and Rowell, 1986).

The experiment presented here is significantly different. It was designed from the beginning such that the stresses generated during loading would be confined to the back portion of the slab and would not extend to the front surface, from where the IR emission was to be measured. Because of the large distance between this surface and the stressed rock volume, >40 cm, we can rule out that frictional heat produced in the stressed rock volume would have diffused to the emitting surface within a time as short as 40 s. Hence, whatever changes we would observe, we can be sure that they are not caused by a flow of Joule heat to the emitting surface. Instead, we have to consider other processes capable of producing rapid changes in the IR emission characteristics upon loading the rock.

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

The most likely candidates are highly mobile electronic charge carriers that have been shown to exist in igneous rocks such as granite, anorthosite and gabbro, albeit in a dormant, electrically inactive form (Freund, 2002; Freund et al., 2004a; 2006). In the following we briefly review what these charge carriers are and what is known about them.

It is widely assumed in the geoscience community that oxygen anions in naturally occurring minerals are fixed in their 2- valence state. However, it has been shown some time ago that a reaction exists by which oxygen anions can change their valence from 2- to 1- (Martens et al., 1976). The reaction involves hydroxyl pairs that are ubiquitous in all minerals, which crystallize in the presence of H₂O, even those that are nominally anhydrous. The hydroxyl pairs in the matrix of these nominally anhydrous minerals can split off H₂ while simultaneously converting their O²⁻ to the 1- (peroxy) state.

A single O⁻ in a matrix of O²⁻ represents a defect electron, resident in the valence band of otherwise insulating materials, also known as a positive hole or p-hole for short. Normally O⁻ form pairs, which we call positive hole pairs, PHP. In oxide materials the PHPs form peroxy anions, O₂²⁻ (Batllo et al., 1991; Freund et al., 1993). In silicate minerals they form peroxy links, O₃X^{OO}XO₃ with X=Si⁴⁺, Al³⁺ etc., well-known in fused silica (Ricci et al., 2001) but believed to be prevalent in all minerals that have crystallized in the presence of H₂O, incorporating some “water” in the form of hydroxyl and thus acquiring the capacity to convert hydroxyl pairs to peroxy links (Freund, 2003). This makes all igneous and high-grade metamorphic rocks candidates for containing PHPs.

As long as the O⁻ exist in the PHP or peroxy state, the p-holes are self-trapped and, hence, electrically inactive. When deviatoric stresses are applied, the rocks begin to plastically deform through dislocations moving in large numbers through the mineral grains (Miguel et al., 2001; Moore and Lockner, 1995; Ohnaka, 1995). When moving dislocations intersect the PHPs, they instantly cause them to break apart and to release p-hole charge carriers (Freund et al., 2006).

As defect electrons in the valence band of the otherwise insulating silicate mineral,

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

the p-holes can move on the O 2sp-dominated levels close to the upper edge of the valence band by exchanging an electron with a neighboring O^{2-} . Theoretically, the speed with which the p-holes could travel is on the order of 300 m/s, consistent with the group velocities in the range of 100–300 m/s measured experimentally (Freund, 2002). Because the valence band forms an energetic continuum, the p-holes can jump grain boundaries. In experiments such as described here where we stress a portion of a large slab of rock the p-holes can flow out of the stressed volume and into unstressed rock (Freund et al., 2006). When they reach the surface, they form a positive charge layer (Freund et al., 1993; King and Freund, 1984). Typical surface potentials measured under open circuit conditions are +1.5 V to +1.75 V. The corresponding charge densities are 10^{-5} Coulomb/m² or 10^{13} – 10^{14} p-holes/m² (Takeuchi et al., 2006; Takeuchi and Nagahama, 2002b). These values give an estimate, by order of magnitude, of the number of p-hole charge carriers that arrive at the rock surface.

Of special interest to the IR emission is the fact that it costs energy to break a peroxy bond or PHP. This energy is spent during the mechanical work that is put into the system during plastic deformation. If the p-holes recombine at the surface to reconstitute the peroxy bond or PHP, some of this energy will be regained. In the primary step this recombination energy will be deposited into the newly formed $O^- - O^-$ bonds, causing them to be “born” in a vibrationally highly excited state. In order to estimate the energies involved we need to know, at least approximately, the O-O bond energy.

In order to photodissociate a peroxy link in fused silica a relatively high energy is needed, 9.4 eV (Nishikawa et al., 1990). From measurements of the electrical conductivity as a function of temperature, the energy needed to thermally dissociate the peroxy bond and release p-hole charge carriers is much lower. In MgO single crystals the activation energy is on the order of 2.4 eV (Freund et al., 1993). We expect similar activation energies for the break-up of peroxy links in silicate minerals and, by extension, in rocks (Freund, 2003). Furthermore it is probably correct to expect that the energy needed for the break-up of PHPs during mechanical deformation is not drastically different. Hence, we can adopt a value around 2.4 eV as being maximally

available during recombination of p-holes to form PHPs.

In Fig. 7 we represent the surface by three corner-linked SiO_4 tetrahedra, two of which terminate at the surface with non-bonded oxygens. On the left, two p-holes are shown to arrive at the surface and to settle on these two adjacent oxygen anions changing them from O^{2-} to O^- . On the right, the two O^- are shown to snap together to form the short O^- - O^- bond characteristic of the peroxy link or PHP (Ricci et al., 2001).

If approximately 2.4 eV are released during p-hole recombination, the new O^- - O^- bond will be “born” in a vibrationally highly excited state. To dissipate this excess energy two decay channels are available: (i) a radiative decay by emitting photons at the characteristic energies of the O-O vibrational manifold or (ii) a non-radiative decay by channeling energy into neighboring Si-O and Al-O bonds, which in turn become excited and emit at their characteristic vibrational frequencies.

The 300 K emission spectrum in Fig. 3 arises from downward transitions of vibrational modes that are thermally excited at 300 K. The probability to populate the excited levels E_n above the ground state E_0 is given by $\exp[-(E_n-E_0)/kT]$, where k is the Boltzmann constant and T the absolute temperature. We are interested in the IR region around 1000 cm^{-1} or $10\text{ }\mu\text{m}$, i.e. in energy levels separated by $\approx 100\text{ meV}$. At 300 K, the mean thermal energy, $kT_{300\text{K}}$, is $\approx 25\text{ meV}$. Hence, the probability to populate the first excited level $n=1$, is $e^{-4} \approx 2 \times 10^{-2}$ or $\approx 2\%$. To populate the second excited level, $n=2$, $\approx 200\text{ meV}$ above E_0 , the probability drops to $e^{-8} \approx 10^{-4}$ or $\approx 0.02\%$ etc. Therefore, for the $10\text{--}12\text{ }\mu\text{m}$ region, levels above $n=1$ are sparsely populated at 300 K and nearly all IR intensity emitted at this temperature comes from downward transitions from the $n=1$ level to the $n=0$ ground level. Emission bands arising from transitions between higher excited levels, or instance from $n=2$ to $n=1$ or from $n=3$ to $n=2$, are called “hot bands”.

The difference spectra in Fig. 5 emphasize the narrow emission bands that are observed from the very on-set of loading, in particular three narrow bands on the right. The band at 930 cm^{-1} ($10.75\text{ }\mu\text{m}$) is consistent with the fundamental stretching mode of the O^- - O^- bond, corresponding to an $n=1$ to $n=0$ transition. For the peroxy link $\text{O}_3\text{Si}^{\text{OO}}\backslash\text{SiO}_3$ in amorphous SiO_2 the energy for this mode has been obtained from

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

quantum-mechanical calculations to be $920\text{--}930\text{ cm}^{-1}$ ($10.75\text{--}10.87\text{ }\mu\text{m}$) (Ricci, et al., 2001). The energies of “hot” transitions for the O-O bond, from $n=2$ to $n=1$ and from $n=3$ to $n=2$, are unknown. However, as a result of the asymmetry of the bond potential, these transitions must occur at slightly lower energies, i.e. lower wavenumbers and longer wavelengths. The two narrow bands at 870 cm^{-1} ($11.5\text{ }\mu\text{m}$), and 810 cm^{-1} ($12.35\text{ }\mu\text{m}$) are consistent with O-O “hot” bands arising from the $n=2$ to $n=1$ and $n=3$ to $n=2$ transitions respectively. In other words, these bands arise from a non-thermal emission process.

The intensity evolution of the three bands at 930 cm^{-1} ($10.75\text{ }\mu\text{m}$), 870 cm^{-1} ($11.5\text{ }\mu\text{m}$), and 810 cm^{-1} ($12.35\text{ }\mu\text{m}$) is replotted in Fig. 8 for the first three 2 min intervals after beginning of loading. The bands at 810 cm^{-1} ($12.35\text{ }\mu\text{m}$) and 870 cm^{-1} ($11.5\text{ }\mu\text{m}$), which we assign to hot transitions, exhibit higher intensities in the beginning. Later, the fundamental at 930 cm^{-1} ($10.75\text{ }\mu\text{m}$) gains intensity, while the hot bands lose their narrow character.

The narrow bands in the $1000\text{--}1200\text{ cm}^{-1}$ ($8.3\text{--}10\text{ }\mu\text{m}$) range in Fig. 5 display a similar intensity evolution as a function of time and load. They are consistent with the concept shown in Fig. 7 which depicts (on the right) that the vibrationally highly excited O^-O^- bond channels energy onto its Si-O and Al-O neighbors, thereby exciting their characteristic vibrational modes.

The excess intensity emitted integrated over the entire course, from the beginning of loading to the failure of the rock, exhibits maxima around 1150 and 1300 cm^{-1} (8.7 and $7.7\text{ }\mu\text{m}$ respectively) but minima close to where the 300 K spectrum has its emission maxima. This suggests that the energy gained by p-hole recombination at the surface is primarily used for non-thermal emission processes, but much less to excite the pool of lattice modes, i.e. to “heat” the rock surface. This implies that most of the excess energy of the newly formed O^-O^- bonds is radiated as discrete IR photons at 930 cm^{-1} ($10.75\text{ }\mu\text{m}$), 870 cm^{-1} ($11.5\text{ }\mu\text{m}$), and 810 cm^{-1} ($12.35\text{ }\mu\text{m}$) or used to excite neighboring Si-O and Al-O bonds. As a result, the anorthosite surface does not heat up. The observations presented here therefore support our proposition that the emis-

sion spectrum of the anorthosite rock observed during stressing is non-thermal, arising primarily from stimulated IR luminescence due to p-hole recombination.

To estimate how much energy can be radiated from the surface during p-hole recombination, we need to know (i) how many p-holes arrive at the surface and (ii) how much energy is released per p-hole recombination event.

(i) For a flat surface across a dielectric contrast from $\epsilon=10$ to $\epsilon=1$, theory predicts surface potential on the order of +0.4 V (King and Freund, 1984). Experimentally, steady state surface potentials in the range from +0.1 V to +1.5 V have been recorded (Freund, 2002; Freund, et al., 2006; Takeuchi, et al., 2006; Takeuchi and Nagahama, 2002a), though pulse-like values as high as +12 V to +17 V have been observed during cracking (Enomoto, et al., 1993; Freund, et al., 2004a). A surface potential of +1 V requires a charge density of $\approx 10^{-5}$ Coulomb m^{-2} , equivalent to $\approx 10^{13}$ m^{-2} (Takeuchi and Nagahama, 2002b).

(ii) We can estimate how many p-hole recombination events are needed to produce a 0.15 K increase of radiation temperature. To break a peroxy bond an activation energy around 2.4 eV is needed (Freund, 2003). If this energy were to be fully regained during p-hole recombination, each recombination event would release 2.4 eV, equivalent in terms of thermal energy kT to $\approx 28\,000$ K. In this case, 2×10^5 recombination events over a surface area of 20 cm^2 would suffice to increase its radiative temperature by 0.15 K.

A surface of 20 cm^2 contains $\approx 10^{16}$ O^{2-} anions. Taking the charge carrier density derived above, $\approx 10^{-5}$ Coulomb m^{-2} or $\approx 10^{13}$ charges m^{-2} (Takeuchi and Nagahama, 2002b), a surface of 20 cm^2 can achieve 1 V surface potential by acquiring $\approx 10^{10}$ p-holes. However, assuming full recovery of the activation energy dispensed during the breaking of the peroxy bonds, only 2×10^5 recombination events per unit time would be needed to increase the radiative temperature by 0.15 K. This suggests that only a small fraction of the p-holes, which reach the surface, need to participate at any time in the IR emission process.

Such an order-of-magnitude estimate gives us confidence that the proposed mech-

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

anism for the observed increase in radiative temperature lies within the limit of our basic assumptions. We note that, in our experiment, the rock surface, which emits the IR radiation, was smooth. If it had been rougher by a geometric factor of about 10, the radiative IR flux would probably have increased by the same factor to an equivalent temperature around 1.5 K, closer to the 2–4°C apparent temperature increases reported for the “thermal anomalies” before major earthquakes. Further experiments will be needed to address this and many other questions raised by this study.

5 Conclusions

When we apply stress 40 m cm away from the emitting rock surface and observe rapid changes in the IR intensity and spectrum, within 40 s or less, we can say with some degree of confidence that this is consistent with the flow of p-hole charge carriers from the stressed rock volume to the emitting surface. The rock surface that emits the IR radiation is clearly too far from the stressed rock volume, over 40 cm, for frictional heat to reach the surface within such a short time. The IR emission is consistent with an exothermal recombination of the p-holes to form positive hole pairs, PHP. It is also consistent with a radiative decay of the vibrationally excited O-O bonds and with the channeling of some of their excess energy into neighboring Si-O and Al-O bonds.

It is still too early to say whether or not the “thermal anomalies” seen in night-time satellite images prior to major seismic activity are caused by p-holes recombining at the surface of the Earth. The experiment described here and its interpretation, however, suggest that p-hole recombination provides a physically plausible mechanism, which could explain the increased IR flux from the surface of a rock subjected to deviatoric stress far from the emitting surface.

The flow of p-holes from the stressed rock volume to the surface as proposed here represents a new mechanism for transporting energy rapidly over macroscopic distances.

Acknowledgements. This work would not have been possible without support from Zheng-

Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

ming Wan and his coworkers Y. Zhang and Q. Zhang, ICESSE, University of California at Santa Barbara, who brought their BOMEN radiospectrometer and other radiometric equipment three times to our laboratory. They helped us perform the experiments but do not necessarily subscribe to our interpretation of the results. We thank M. Jhabvala, NASA Goddard Space Flight Center, for encouraging us and helping us during the early phase of this project. We thank G. Pacchioni, Milano, Italy for performing additional calculations for the O-O vibrational energy in the feldspar structure. We acknowledge financial support by grants from the NASA Ames Research Center Director's Discretionary Fund. B.W.S.L. was supported by a grant from the National Geospatial Agency (NGA). A.T. acknowledges support from the Japan Society for the Promotion of Science (JSPS) for Young Scientists.

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Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

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Stimulated infrared emission from rocks

F. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

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Stimulated infrared emission from rocksF. Freund et al.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

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Stimulated infrared emission from rocksF. Freund et al.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

I◀

▶I

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

Stimulated infrared emission from rocks

F. Freund et al.

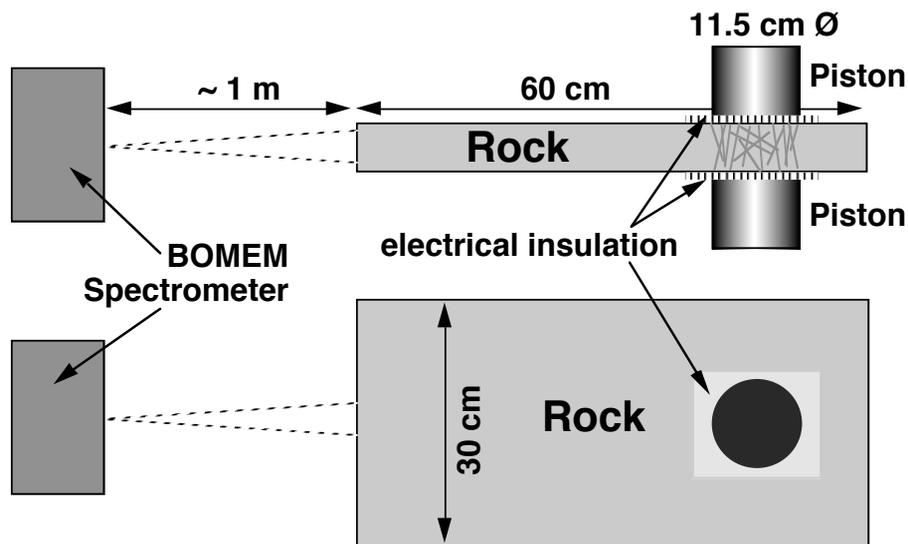


Fig. 1. Schematic of the set-up used to measure the IR emission from the flat front face of a $60 \times 30 \times 7.5 \text{ cm}^3$ block of anorthosite, loaded over 40 cm away from the emitting rock surface.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

Stimulated infrared emission from rocks

F. Freund et al.

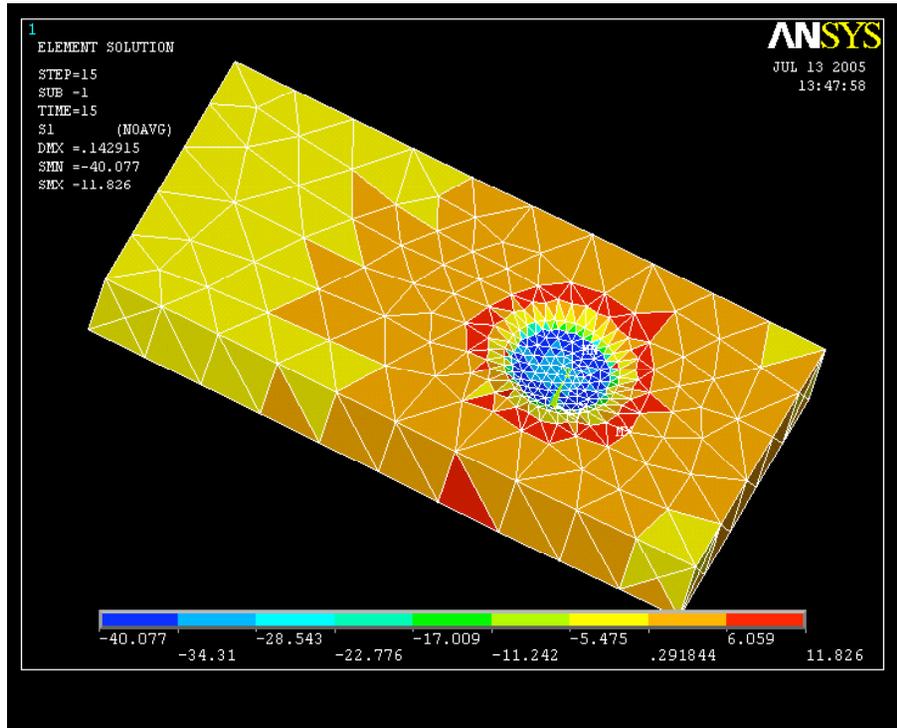


Fig. 2. Finite analysis representation, using a variable grid size, of the stress and strain distribution in the anorthosite block during asymmetric loading. The 20 cm^2 surface from where the IR emission is measured is on the hidden surface on the left.

Title Page

Abstract

Introduction

Conclusions

References

Tables

Figures

◀

▶

◀

▶

Back

Close

Full Screen / Esc

Printer-friendly Version

Interactive Discussion

**Stimulated infrared
emission from rocks**

F. Freund et al.

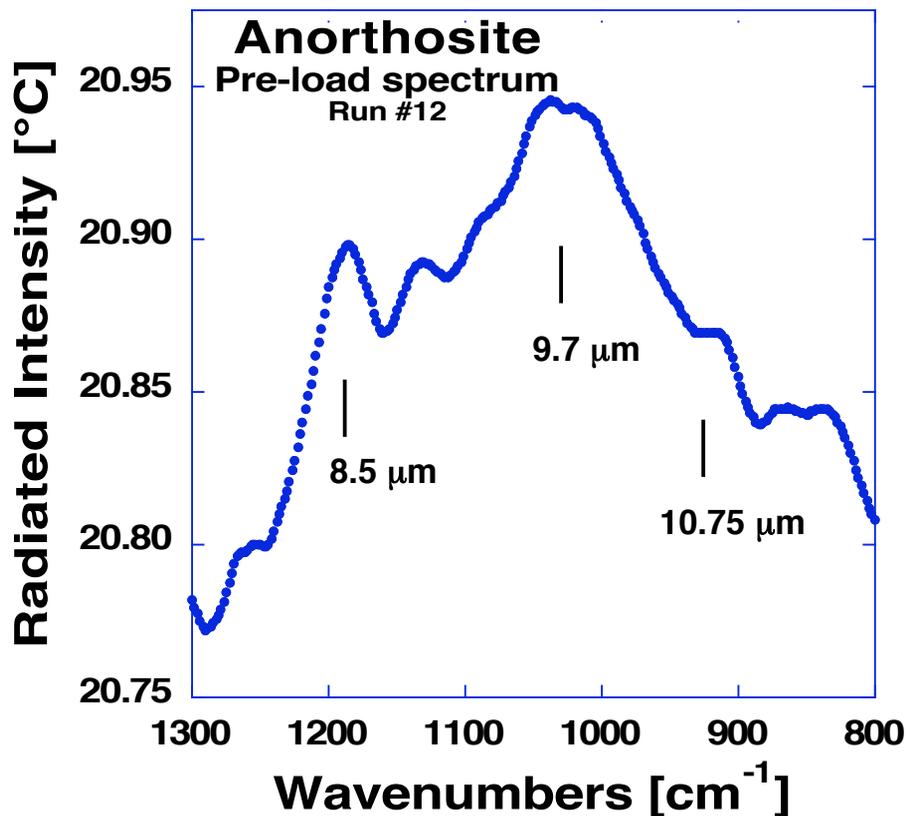


Fig. 3. IR emission spectrum at room temperature (average of 10 files of 25 scans each) from the flat front surface of the anorthosite block before loading.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

Stimulated infrared emission from rocks

F. Freund et al.

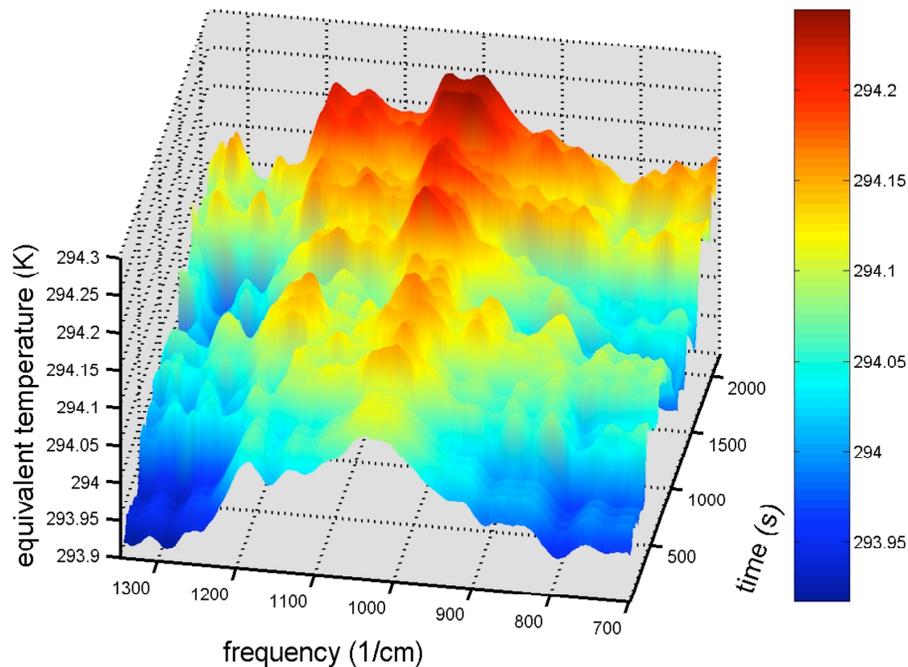


Fig. 4. 3-D plot of the intensity evolution and spectral changes of the IR emission between $700\text{--}1350\text{ cm}^{-1}$ ($7.37\text{--}14.3\text{ }\mu\text{m}$) from the front face of the anorthosite block before and during loading up to failure, plotted as a function of time during loading (for details see text).

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

**Stimulated infrared
emission from rocks**

F. Freund et al.

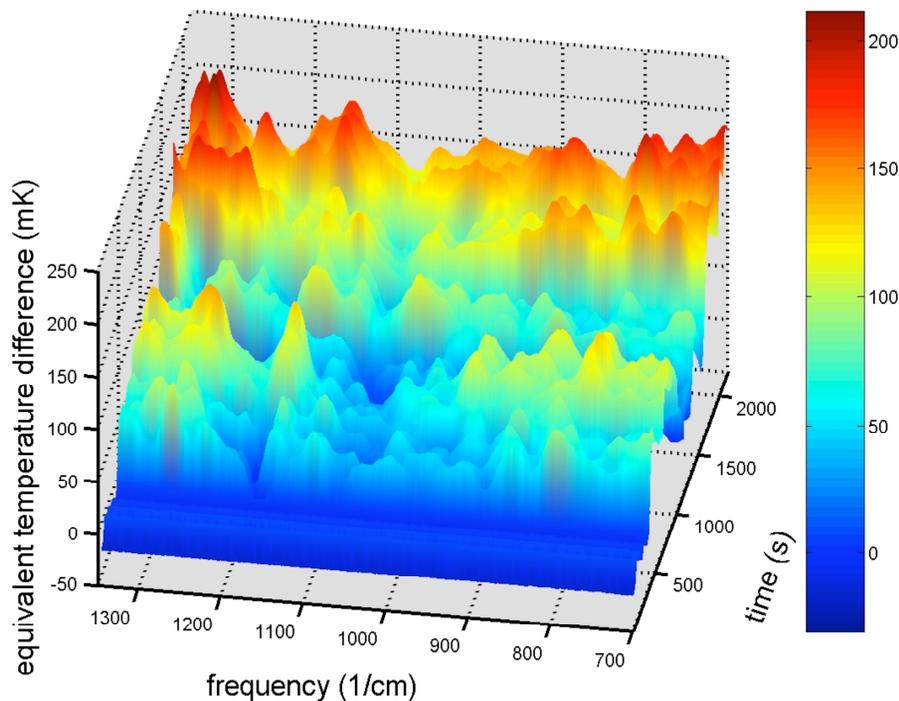


Fig. 5. Difference plot of the intensity evolution and spectral changes of the IR emission from the front face of the anorthosite block obtained by subtracting each file recorded during loading from the average of the pre-load files.

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

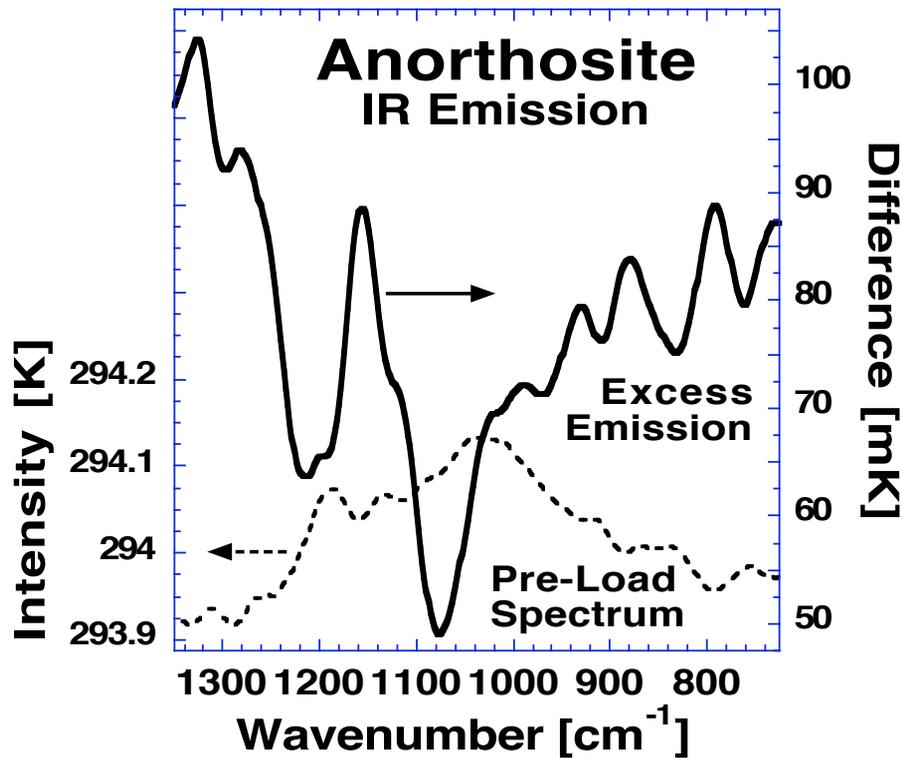


Fig. 6. Total excess intensity emitted over the 7.37–14.7 μm range from the front face of the anorthosite block during loading (solid line) compared to the pre-load spectrum (dotted line).

[Title Page](#)
[Abstract](#)
[Introduction](#)
[Conclusions](#)
[References](#)
[Tables](#)
[Figures](#)
[◀](#)
[▶](#)
[◀](#)
[▶](#)
[Back](#)
[Close](#)
[Full Screen / Esc](#)
[Printer-friendly Version](#)
[Interactive Discussion](#)

**Stimulated infrared
emission from rocks**

F. Freund et al.

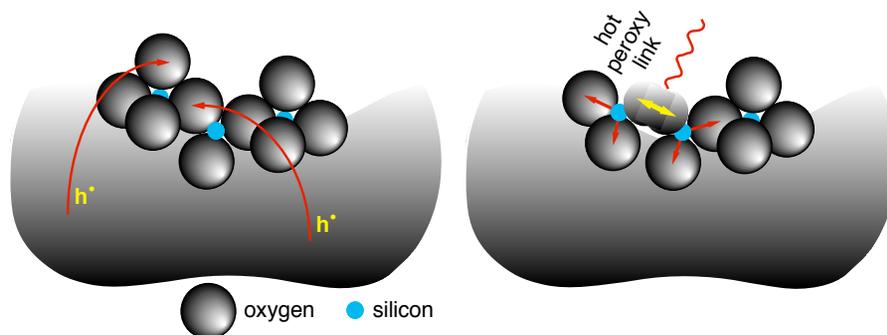


Fig. 7. Schematic representation of a mineral surface to illustrate the processes that may take place when p-holes arrive at the surface (left). When the two p-holes recombine, recombination energy is released, leading to a vibrationally highly excited O-O bond, which can de-excite radiatively by emitting IR photons characteristic of transitions the energy levels of the O-O bond, and non-radiatively by channeling energy into neighboring bonds (right).

[Title Page](#)[Abstract](#)[Introduction](#)[Conclusions](#)[References](#)[Tables](#)[Figures](#)[◀](#)[▶](#)[◀](#)[▶](#)[Back](#)[Close](#)[Full Screen / Esc](#)[Printer-friendly Version](#)[Interactive Discussion](#)

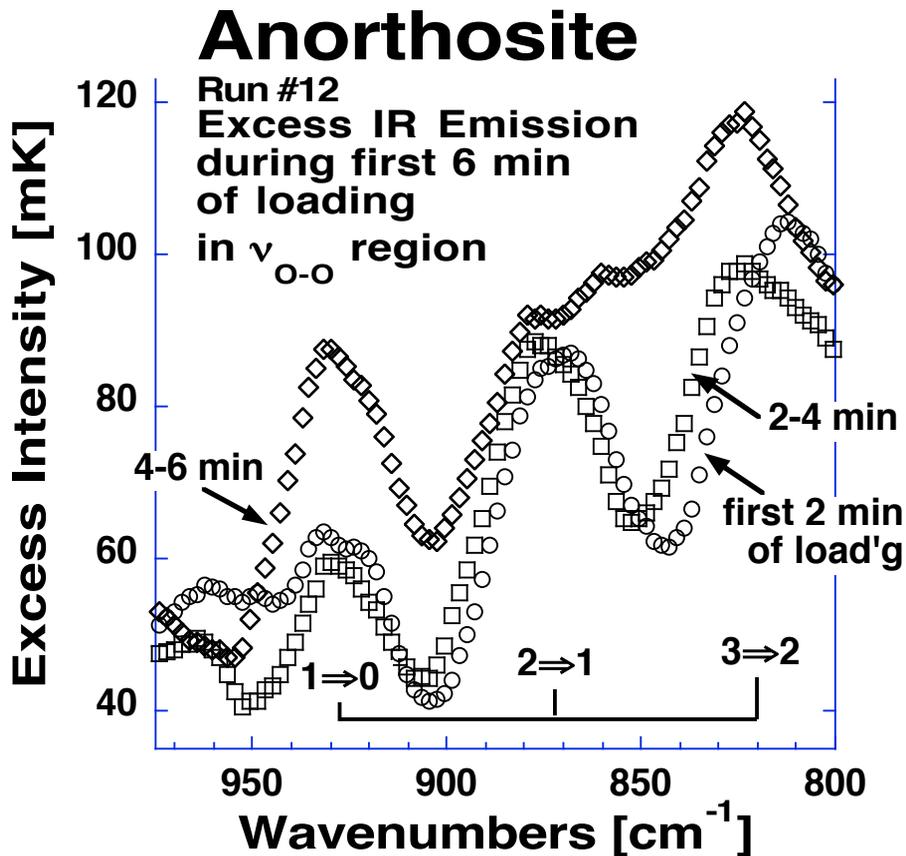


Fig. 8. Evolution of the IR emission bands in the spectral range expected to contain the O-O fundamental and “hot bands” during the first 6 min of loading, broken down in 2 min intervals.

[Title Page](#)
[Abstract](#)
[Introduction](#)
[Conclusions](#)
[References](#)
[Tables](#)
[Figures](#)
[◀](#)
[▶](#)
[◀](#)
[▶](#)
[Back](#)
[Close](#)
[Full Screen / Esc](#)
[Printer-friendly Version](#)
[Interactive Discussion](#)